

# MUS421 Lecture 3A

## FFT Windows

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### Outline

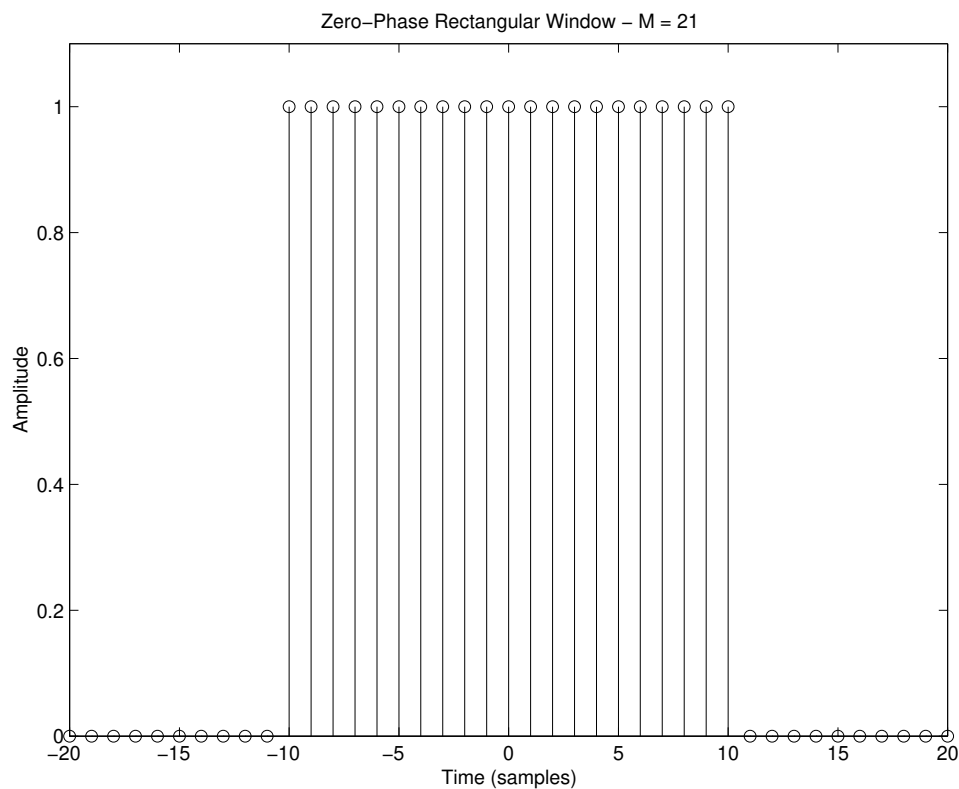
- Rectangular, Hann, Hamming
- MLT Sine
- Blackman-Harris Window Family
- Bartlett
- Poisson
- Slepian and Kaiser
- Dolph-Chebyshev
- Gaussian
- Optimal Windows

# The Rectangular Window

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Previously, we looked at the rectangular window:

$$w_R(n) \triangleq \begin{cases} 1, & |n| \leq \frac{M-1}{2} \\ 0, & \text{otherwise} \end{cases}$$



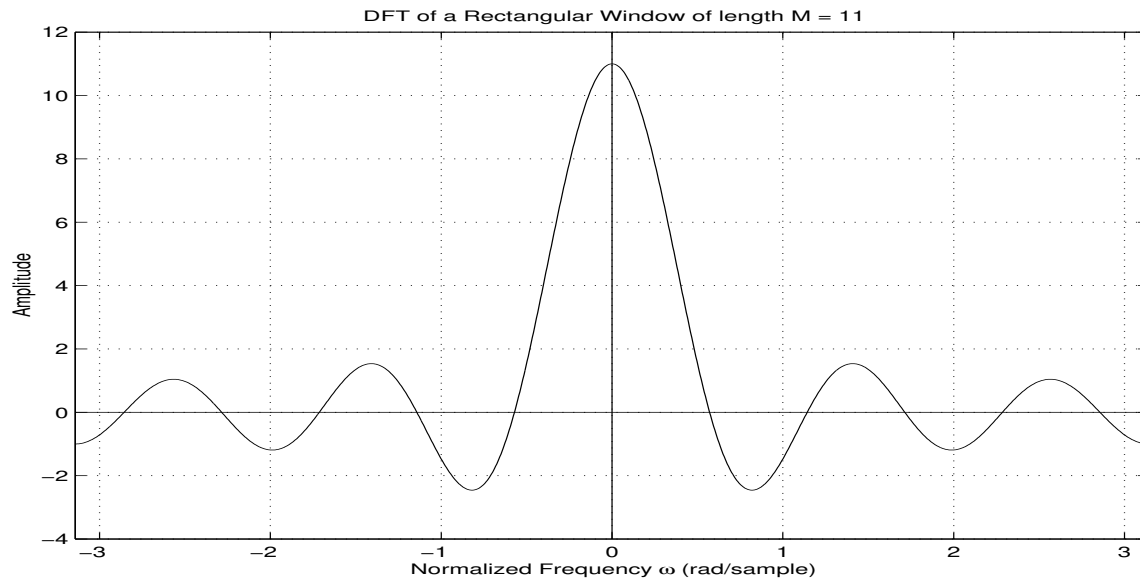
The window transform (DTFT) was found to be

$$W_R(\omega) = \frac{\sin\left(M\frac{\omega}{2}\right)}{\sin\left(\frac{\omega}{2}\right)} \triangleq M \cdot \text{asinc}_M(\omega) \quad (1)$$

where  $\text{asinc}_M(\omega)$  denotes the *aliased sinc function*.

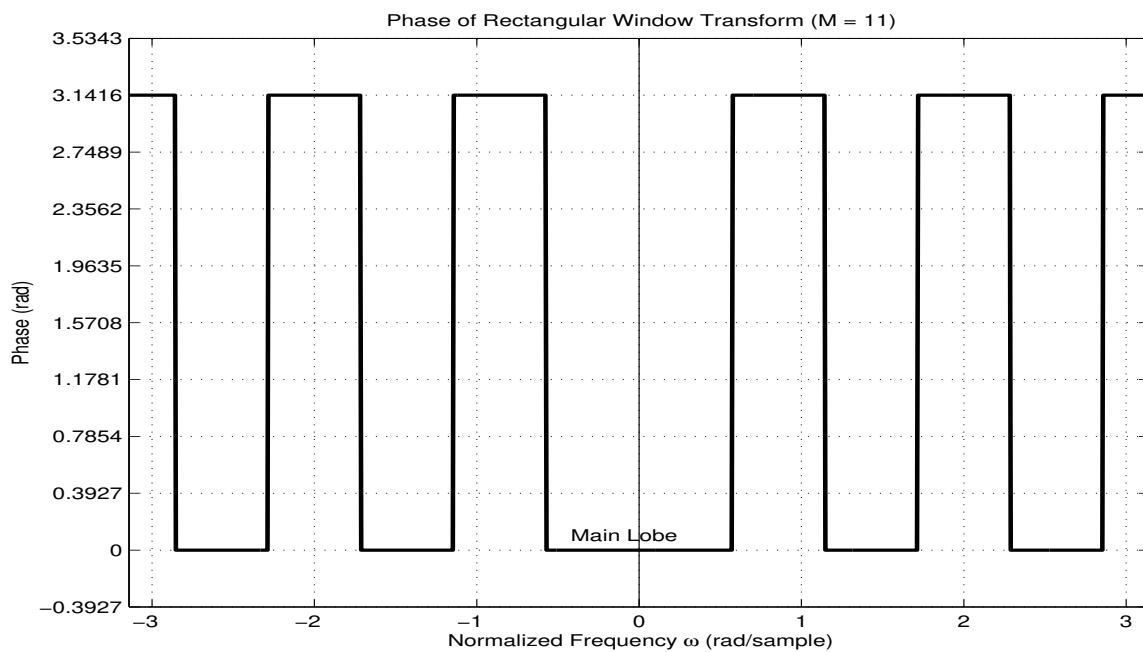
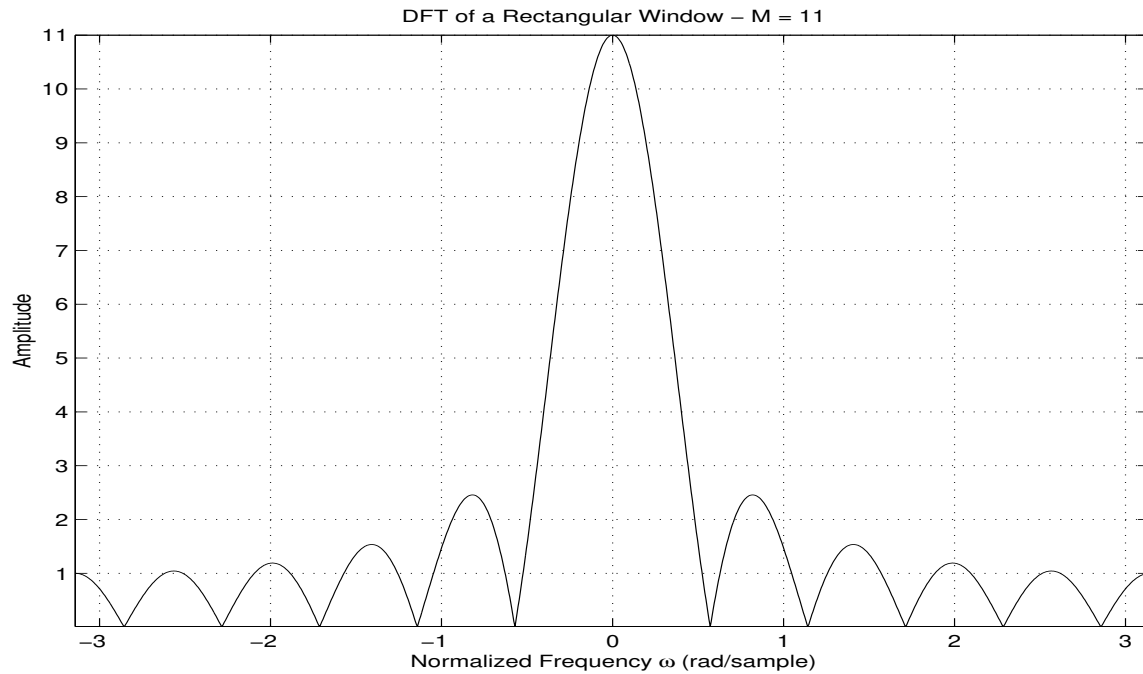
$$\text{asinc}_M(\omega) \triangleq \frac{\sin(M\omega/2)}{M \cdot \sin(\omega/2)}$$

This result is plotted below:

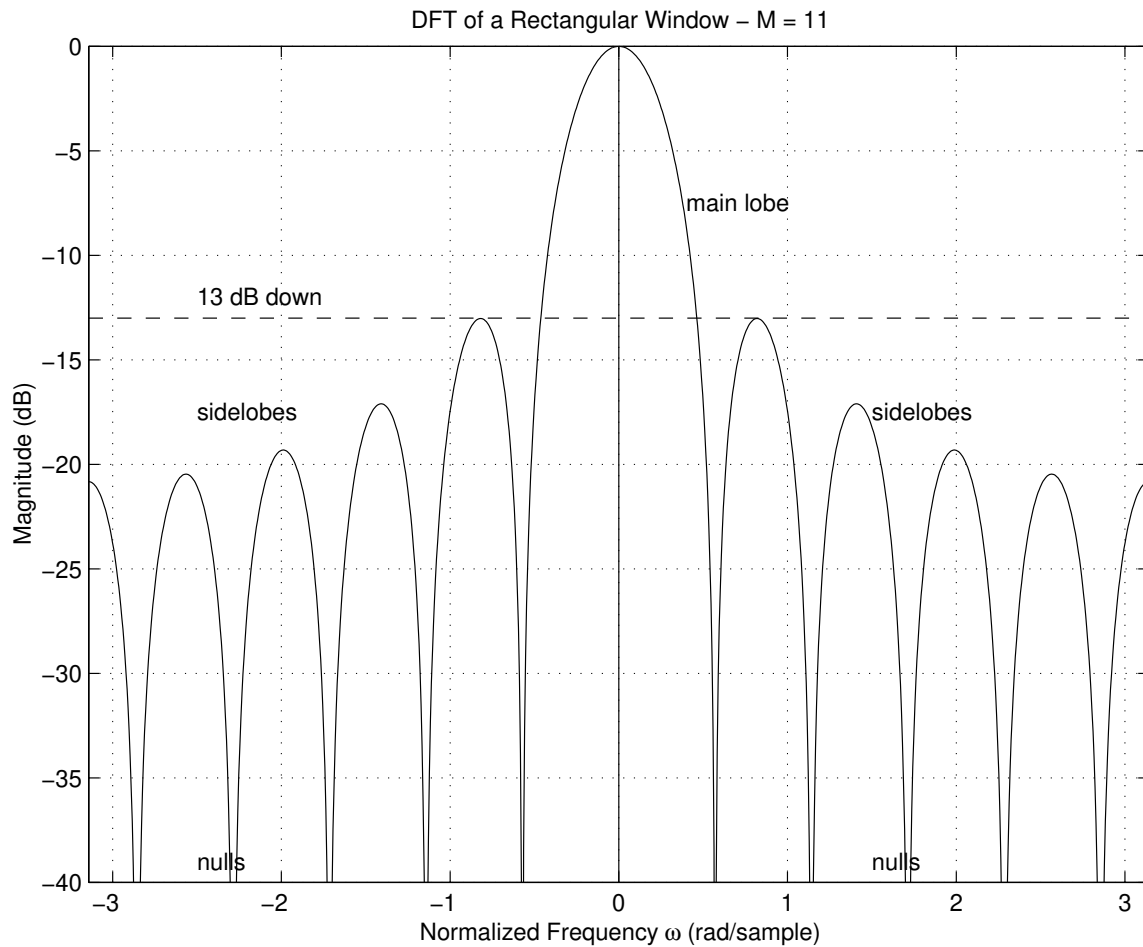


Note that this is the complete window transform, not just its magnitude. We obtain real window transforms like this only for symmetric, zero-centered windows.

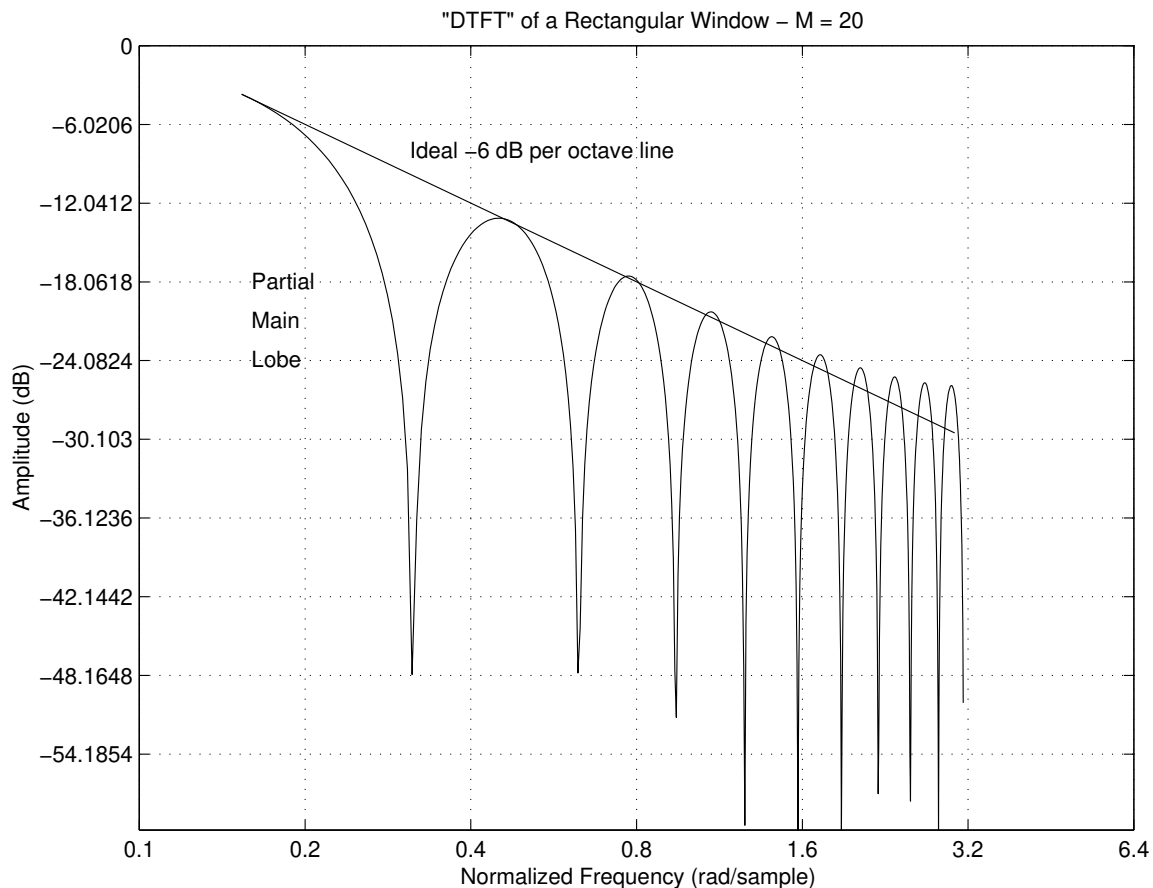
More generally, we may plot both the *magnitude* and *phase* of the window versus frequency:



In audio work, we more typically plot the window transform magnitude on a *decibel (dB) scale*:



Since the DTFT of the rectangular window approximates the sinc function, it should “roll off” at approximately 6 dB per octave, as verified in the log-log plot below:



As the sampling rate approaches infinity, the rectangular window transform converges exactly to the sinc function. Therefore, the departure of the roll-off from that of the sinc function can be ascribed to *aliasing* in the frequency domain, due to sampling in the time domain.

## Sidelobe Roll-Off Rate

In general, when only the first  $n$  terms *exist* in the power-series expansion of a continuous function  $w(t)$  (*i.e.*, each term is finite), then the Fourier Transform magnitude  $|W(\omega)|$  is asymptotically proportional to

$$|W(\omega)| \rightarrow \frac{1}{\omega^n} \quad (\text{as } \omega \rightarrow \infty)$$

**Proof:** Papoulis, **Signal Analysis**, McGraw-Hill, 1977

Thus, we have the following rule-of-thumb:

$n \text{ terms} \leftrightarrow -6n \text{ dB per octave roll-off rate}$

(since  $-20 \log_{10}(2) = 6.0205999 \dots$ ).

This is also  $-20n$  dB per *decade*.

To apply this result, we normally only need to look at the window's *endpoints*. The interior of the window is usually differentiable of all orders.

### Example Roll-Off Rates:

- Amplitude discontinuity ( $n = 1$ )  $\leftrightarrow -6$  dB/octave
- Slope discontinuity ( $n = 2$ )  $\leftrightarrow -12$  dB/octave
- Curvature discontinuity ( $n = 3$ )  $\leftrightarrow -18$  dB/octave

For discrete-time windows, the roll-off rate slows down at high frequencies due to aliasing.

In summary, the DTFT of the  $M$ -sample rectangular window is proportional to the ‘aliased sinc function’:

$$\begin{aligned}\text{asinc}_M(\omega T) &\triangleq \frac{\sin(\omega MT/2)}{M \cdot \sin(\omega T/2)} \\ &\approx \frac{\sin(\pi f MT)}{M \pi f T} \triangleq \text{sinc}(f MT)\end{aligned}$$

Points to note:

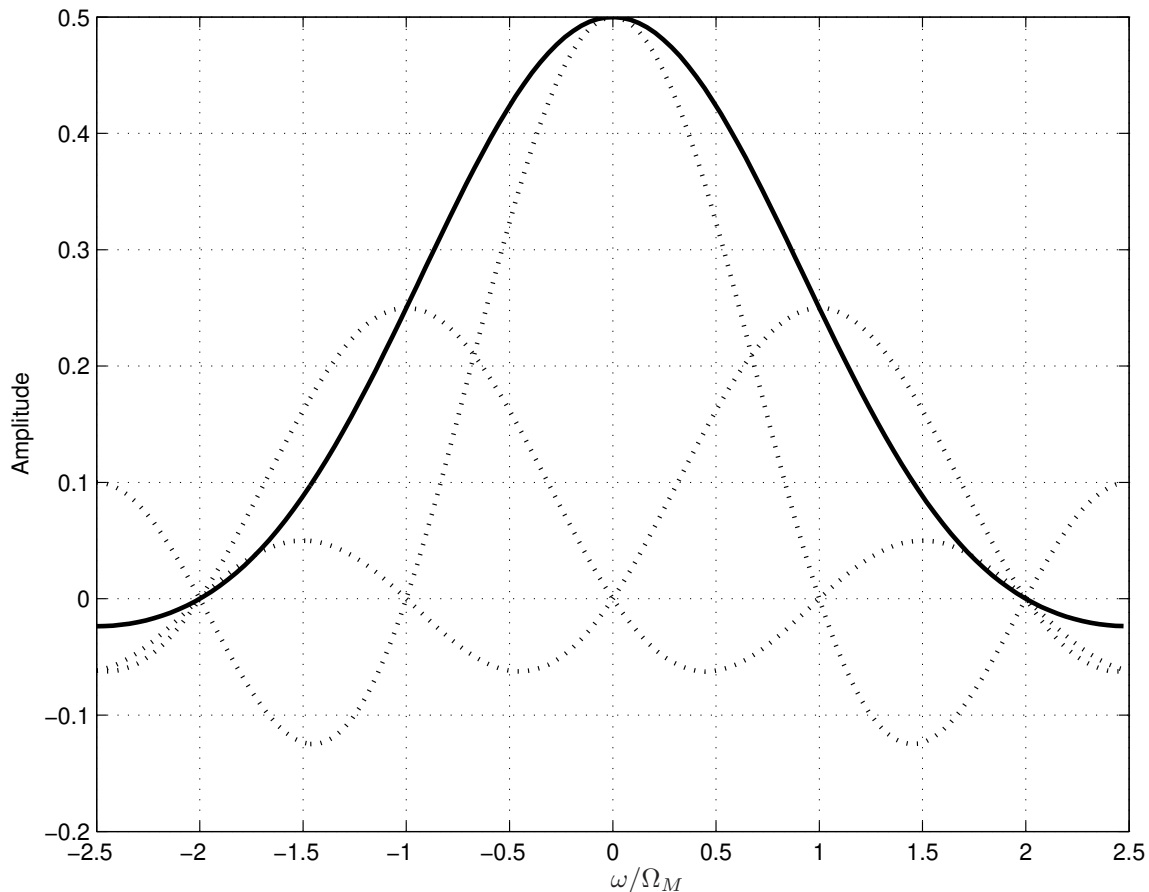
- Zero crossings at integer multiples of  $\Omega_M \triangleq \frac{2\pi}{M}$   
where  $\Omega_M \triangleq \frac{2\pi}{M}$  = frequency sampling interval for a length  $M$  DFT
- Main lobe width is  $2\Omega_M = \frac{4\pi}{M}$
- As  $M$  gets bigger, the mainlobe narrows  
(better frequency resolution)
- $M$  has *no effect on the height of the side lobes*  
(Same as the “Gibbs phenomenon” for Fourier series)
- First sidelobe only 13 dB down from main-lobe peak
- Side lobes roll off at approximately 6 dB per octave
- A *phase term* arises when we shift the window to make it *causal*, while the window transform is real in the zero-centered case (i.e., centered about time 0)



# Generalized Hamming Window Family

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Consider the following picture in the frequency domain:



[https://ccrma.stanford.edu/~jos/Windows/Generalized\\_Hamming\\_Window\\_Family.html](https://ccrma.stanford.edu/~jos/Windows/Generalized_Hamming_Window_Family.html)

We have added 2 extra aliased sinc functions (shifted), which results in the following behavior:

- There is some cancellation of the side lobes
- The width of the main lobe is doubled

In terms of the rectangular window transform

$W_R(\omega) = M \cdot \text{asinc}_M(\omega)$  (zero-centered, unit-amplitude case), this can be written as:

$$W_H(\omega) \triangleq \alpha W_R(\omega) + \beta W_R(\omega - \Omega_M) + \beta W_R(\omega + \Omega_M)$$

Using the Shift Theorem dual, we can take the inverse transform of the above equation:

$$\begin{aligned} w_H &= \alpha w_R(n) + \beta e^{-j\Omega_M n} w_R(n) + \beta e^{j\Omega_M n} w_R(n) \\ &= w_R(n) \left[ \alpha + 2\beta \cos \left( \frac{2\pi n}{M} \right) \right] \end{aligned}$$

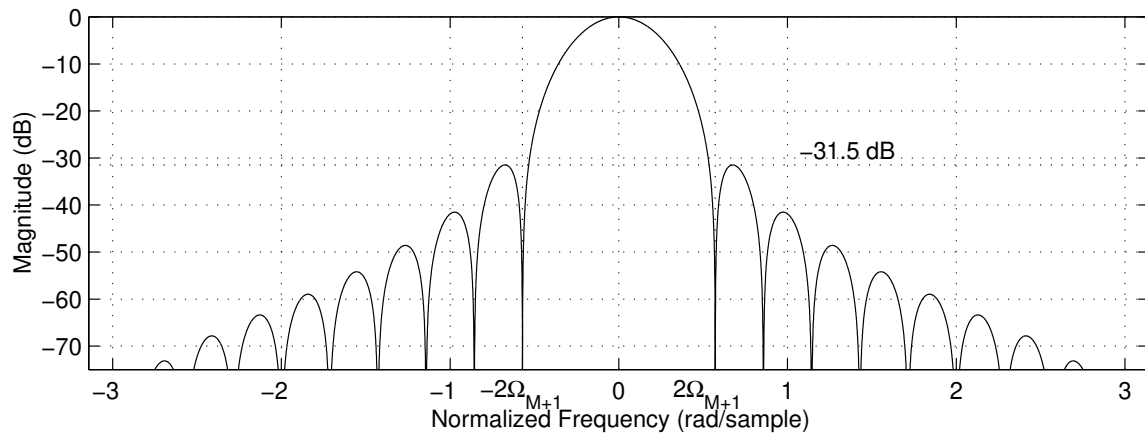
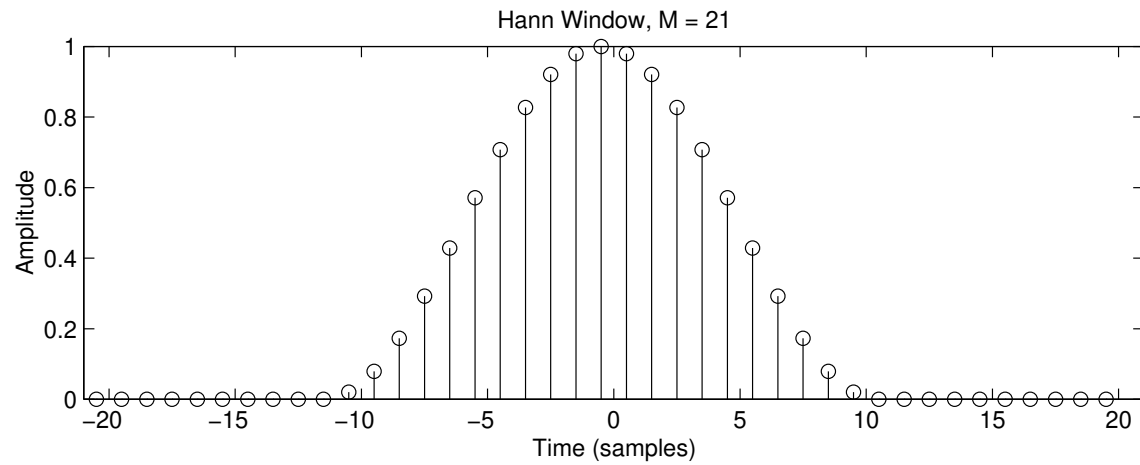
Choosing various parameters for  $\alpha$  and  $\beta$  result in different windows in the generalized Hamming family, some of which have names.

# Hann or Hanning or Raised Cosine

The Hann window is defined by the settings

$\alpha = 1/2$  and  $\beta = 1/4$ :

$$w_H(n) = w_R(n) \left[ \frac{1}{2} + \frac{1}{2} \cos(\Omega_M n) \right] = w_R(n) \cos^2 \left( \frac{\Omega_M}{2} n \right)$$



Hann window properties:

- Main lobe is  $4\Omega_M$  wide
- First side lobe is at  $-31$  dB
- Side-lobes roll off at  $\approx 18$  dB / octave

Compare to the Rectangular window:

- Main lobe is  $2\Omega_M$  wide
- First side lobe at  $-13$  dB
- Side-lobes roll off at  $\approx 6$  dB / octave

## Hamming

This window is determined by choosing  $\alpha$  to cancel the first side lobe and  $\beta$  to normalize peak amplitude to 1 in the time domain:

$$\alpha = \frac{25}{46} \approx 0.54$$

$$\beta = (1 - \alpha)/2$$

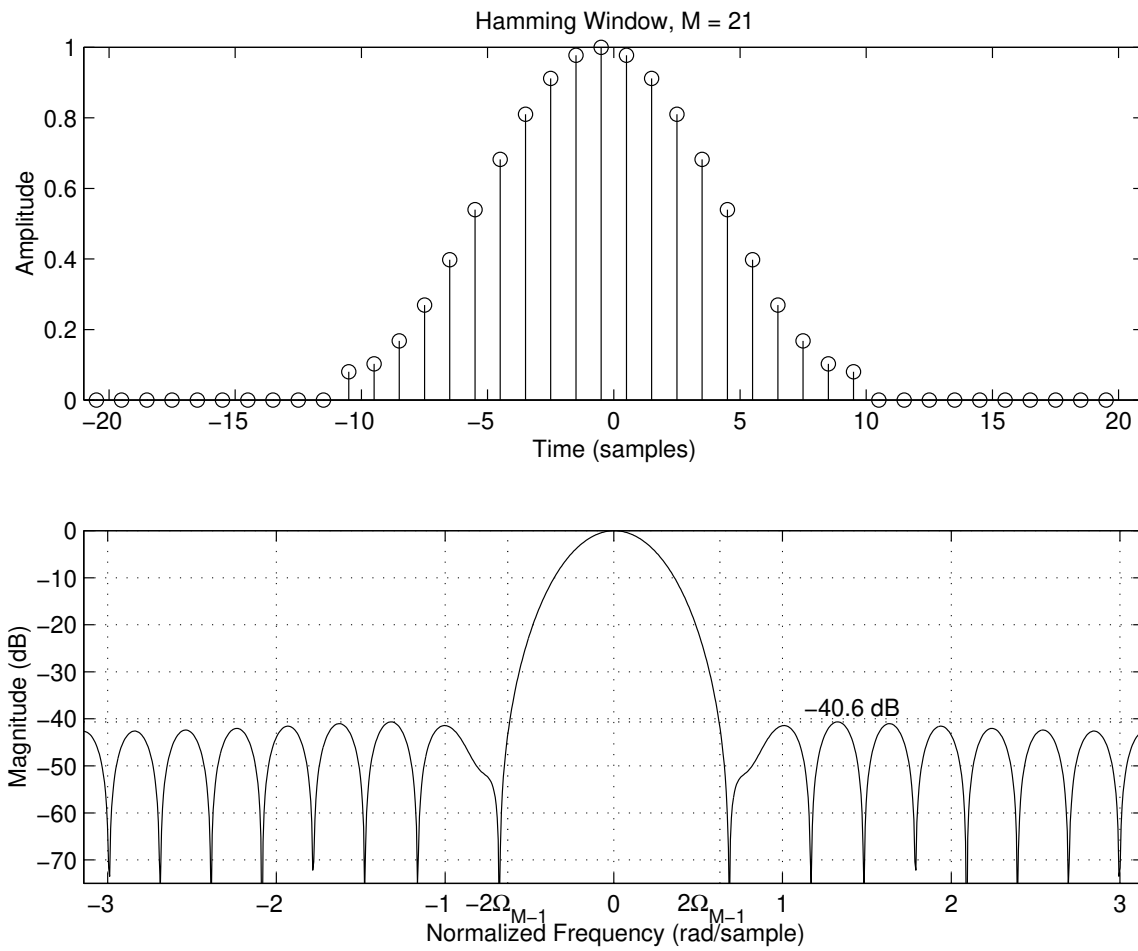
**Note:** The Hamming window is very close to the generalized Hamming window which *minimizes sidelobe level* within the family:

$$\alpha = 0.53836 \quad (\text{minimum peak side-lobe magnitude})$$

Thus, the Hamming window is the “Chebyshev Generalized Hamming Window” rounded to two significant digits.

Chebyshev-type designs generally exhibit *equiripple* error behavior, since the worst-case error (sidelobe level in this case) is minimized (see Dolph-Chebyshev window below)

# Hamming Window

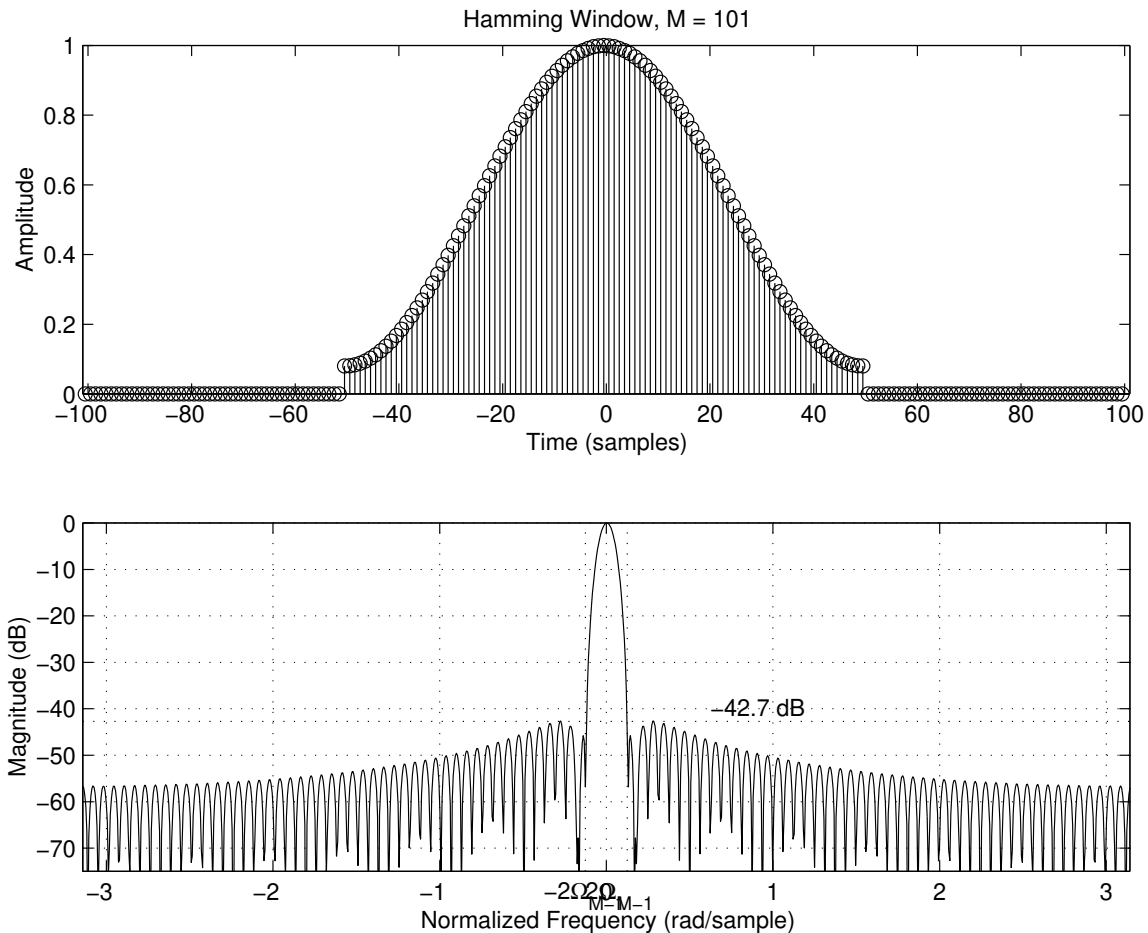


## Hamming Window Properties

- Discontinuous “slam to zero” at endpoints
- main lobe is  $4\Omega_M$  (like Hann)
- Roll off is approx. 6 dB/octave (but aliased)
- 1st side lobe is improved over Hann
- side lobes closer to “equal ripple”

**Question:** How can side-lobes approximate “equal ripple” when they cannot roll-off slower than 6 dB / octave?

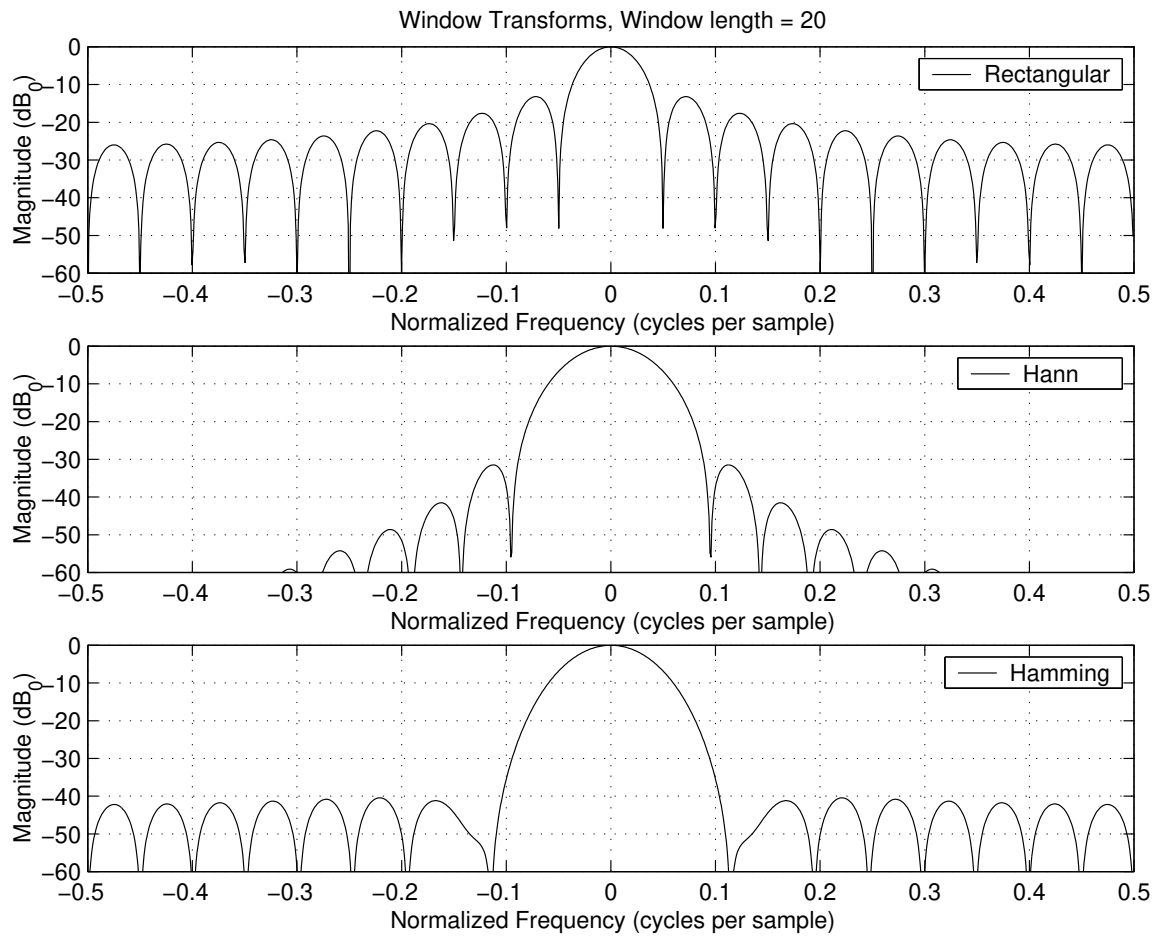
# Longer Hamming Window



- Since the side-lobes nearest the main lobe are most affected by the Hamming optimization, we now have a larger frequency region over which the spectral envelope looks like that of the asinc function (an “aliased -6 dB/octave roll-off”).
- The side-lobe level (-42.7 dB) is also improved over that of the shorter window (-40.6 dB).



# Window Transform Summary



## The MLT Sine Window

The *Modulated Lapped Transform* (MLT) uses the *sine window*:

$$w(n) = \sin \left[ \left( n + \frac{1}{2} \right) \frac{\pi}{2M} \right], \quad n = 0, 1, 2, \dots, 2M - 1.$$

- Used in MPEG-1, Layer 3 (MP3 format), MPEG-2 AAC, MPEG-4
- Sidelobes 24 dB down
- Asymptotically optimal coding gain
- Zero-phase-window transform (“truncated cosine window”) has smallest *moment of inertia* over all windows:

$$\int_{-\pi}^{\pi} \omega^2 W(\omega) d\omega = \min$$

# Blackman-Harris Window Family

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- The Blackman-Harris family of windows is basically a generalization of the Hamming family.
- In the case of the Hamming family, we constructed a summation of 3 shifted sinc functions.
- The Blackman-Harris family is derived by considering a more general summation of shifted sinc functions:

$$w_B(n) = w_R(n) \sum_{l=0}^{L-1} \alpha_l \cos(l\Omega_M n)$$

where  $\Omega_M \triangleq 2\pi/M$ ,  
 $n = -(M-1)/2, \dots, (M-1)/2$ , ( $M$  odd).

## Special Cases:

- $L = 1 \Rightarrow$  Rectangular
- $L = 2 \Rightarrow$  Generalized Hamming
- $L = 3 \Rightarrow$  Blackman Family
- $L > 3 \Rightarrow$  Blackman-Harris Family

## Frequency-Domain Implementation

The Blackman-Harris window family can be very efficiently implemented in the frequency domain as a  $(2L - 1)$ -point convolution with the spectrum of the unwindowed data. Example:

1. Hann Window = 3-Point DFT <sub>$M$</sub>  Smoother:

- Start with a length  $M$  rectangular window
- Take an  $M$ -point DFT
- Convolve the DFT data with the 3-point smoother  $[1/4, 1/2, 1/4]$  to implement a Hann window
- Note that the Hann window requires *no multiplies* in linear fixed-point data formats

2. Any Blackman window is a 5-point smoother for a Length  $M$  (critically sampled) DFT

## Classic Blackman

The so-called “Blackman Window” is the specific case in which  $\alpha_0 = 0.42$ ,  $\alpha_1 = 0.5$ , and  $\alpha_2 = 0.08$

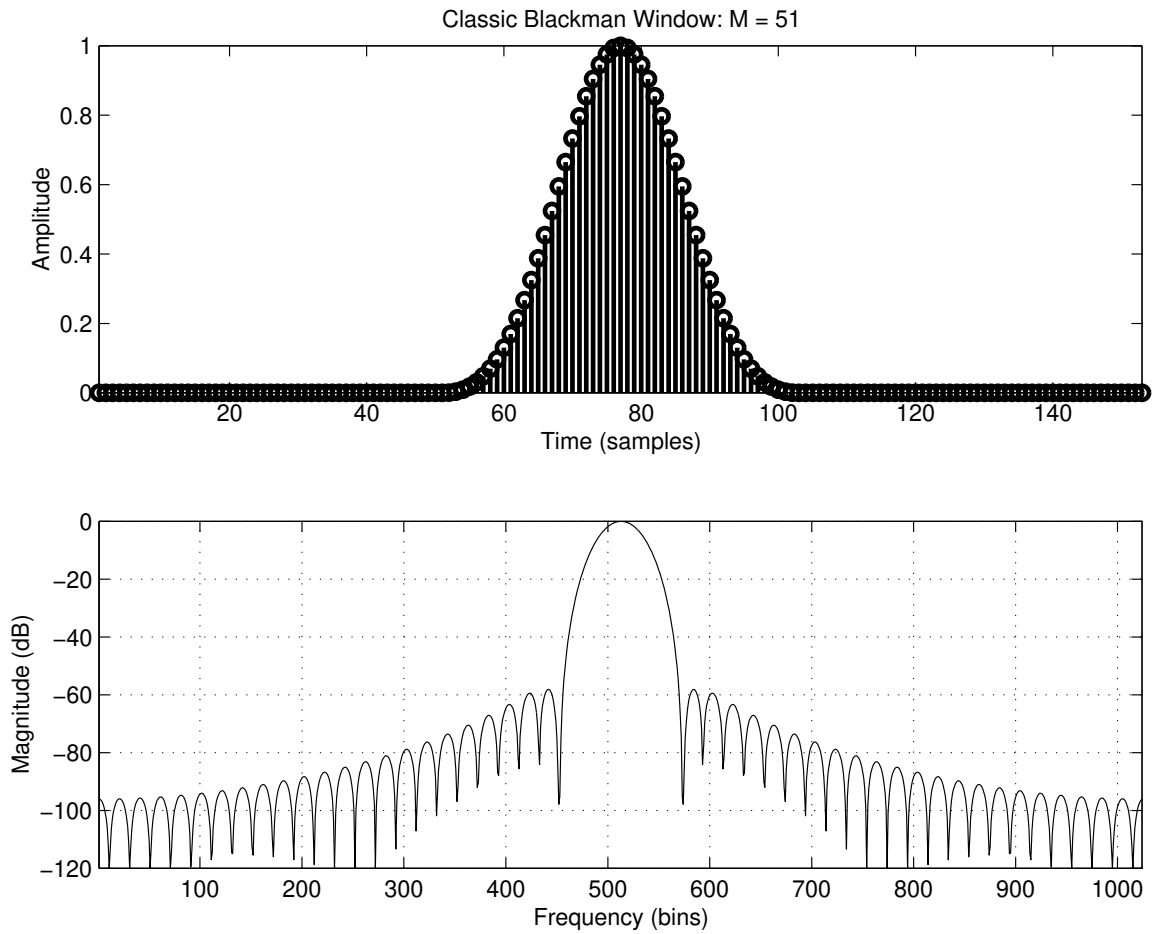
### Properties:

- Sidelobes roll off about 18 dB per octave (as  $T \rightarrow 0$ )
- $-58$  dB sidelobe level (worst case)
- One degree of freedom used to increase the roll-off rate from 6 dB/octave to 18 dB per octave
- One degree of freedom used to minimize sidelobes
- One degree of freedom used to scale the window

### Matlab:

```
N = 101; L = 3; No2 = (N-1)/2; n=-No2:No2;
ws = zeros(L,3*N); z = zeros(1,N);
for l=0:L-1
    ws(l+1,:) = [z,cos(l*2*pi*n/N),z];
end
alpha = [0.42,0.5,0.08]; % Classic Blackman
w = alpha * ws;
```

# Classic Blackman Window and Transform



## Three-Term Blackman-Harris

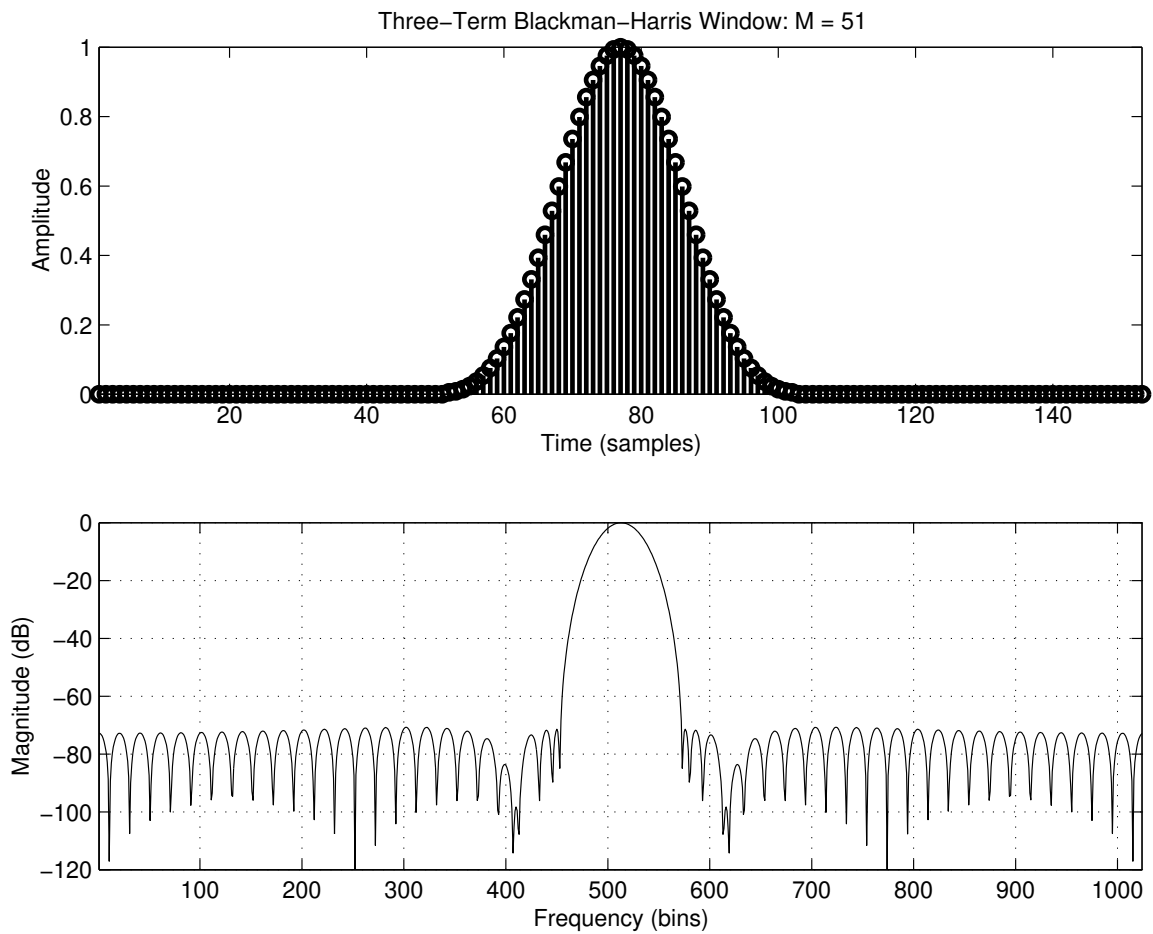
### Properties:

- $\alpha_0 = 0.4243801$   $\alpha_1 = 0.4973406$ , and  $\alpha_2 = 0.0782793$ .
- Side-lobe level  $-71.5$  dB.
- Side lobes roll off  $\approx 6$  dB per octave in the absence of aliasing (like rectangular and Hamming).
- *All degrees of freedom (scaling aside) are used to minimize side lobes* (like Chebyshev-Hamming  $\approx$  Hamming).

### Matlab:

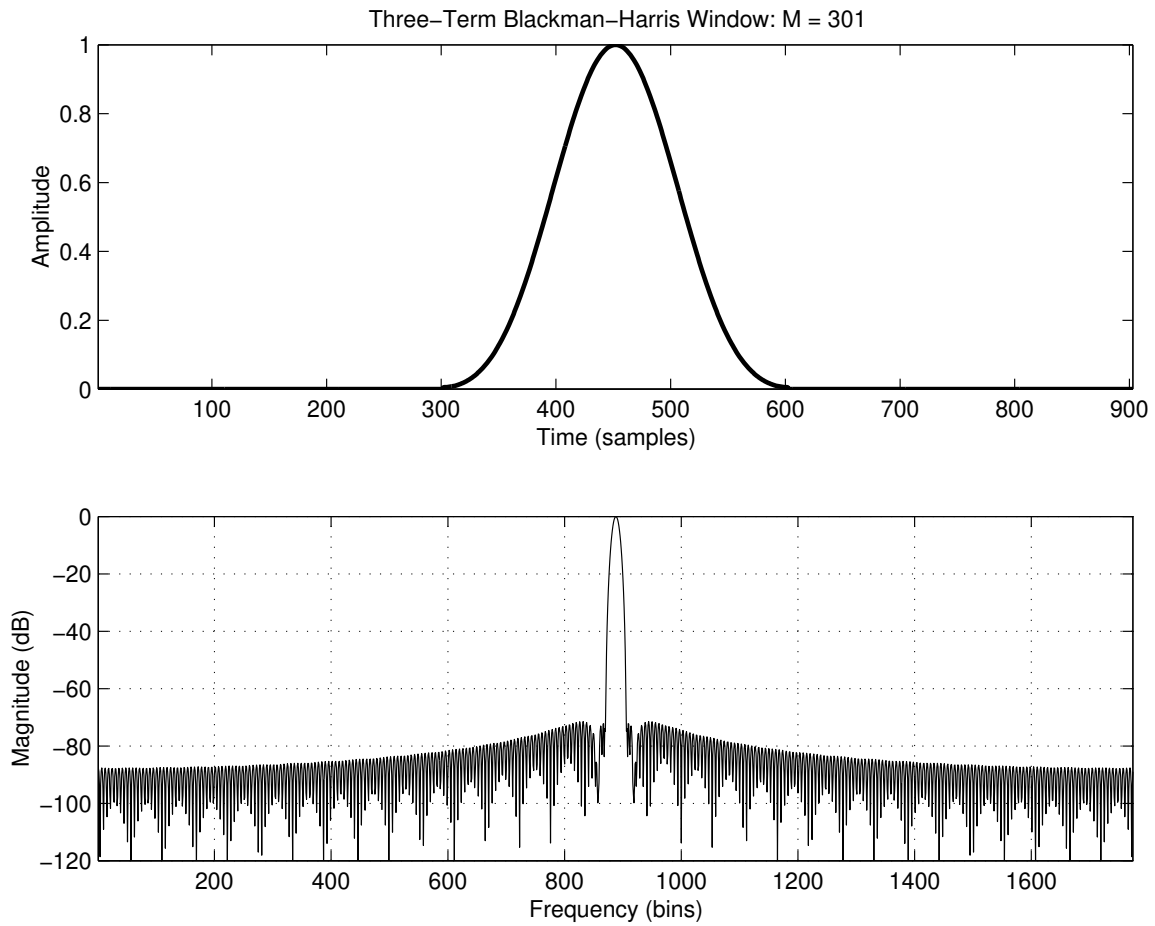
```
N = 101; L = 3; No2 = (N-1)/2; n=-No2:No2;
ws = zeros(L,3*N); z = zeros(1,N);
for l=0:L-1
    ws(l+1,:) = [z,cos(l*2*pi*n/N),z];
end
% 3-term Blackman-Harris(-Nuttall):
alpha = [0.4243801, 0.4973406, 0.0782793];
w = alpha * ws;
```

# Three-Term Blackman-Harris Window and Transform





# Longer Three-Term Blackman-Harris Window and Transform



## Power-of-Cosine

$$w(n) = w_R(n) \cos^P \left( \frac{\pi n}{M} \right), \quad n \in \left[ -\frac{M-1}{2}, \frac{M-1}{2} \right]$$

- $P = 0, 1, 2, \dots$
- first  $P$  terms of its Taylor expansion, evaluated at the endpoints (1/2 sample beyond last sample) are 0
- roll-off rate  $\approx 6(P+1)$  dB/octave
- $P = 0 \Rightarrow$  Rectangular window
- $P = 1 \Rightarrow$  MLT sine window (shifted to zero-phase)
- $P = 2 \Rightarrow$  Hann window (“raised cosine” = “ $\cos^2$ ”)
- $P = 4 \Rightarrow$  Alternate Blackman (max roll-off rate in Blackman family)
- $\dots$

Thus,  $\cos^P$  windows parametrize  $L$ th-order Blackman-Harris windows configured to use all degrees of freedom to maximize roll-off rate ( $L = P/2 + 1$ )

# Spline Windows

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A spline window of order  $N$  is a repeated convolution of rectangular windows:

$$w_{\text{Spline}(N)}(n) = \underbrace{(w_R * w_R * \cdots * w_R)}_{N+1}(n)$$
$$\leftrightarrow W_{\text{Spline}(N)}(\omega) = \text{sinc}^{N+1}$$

## Special Cases:

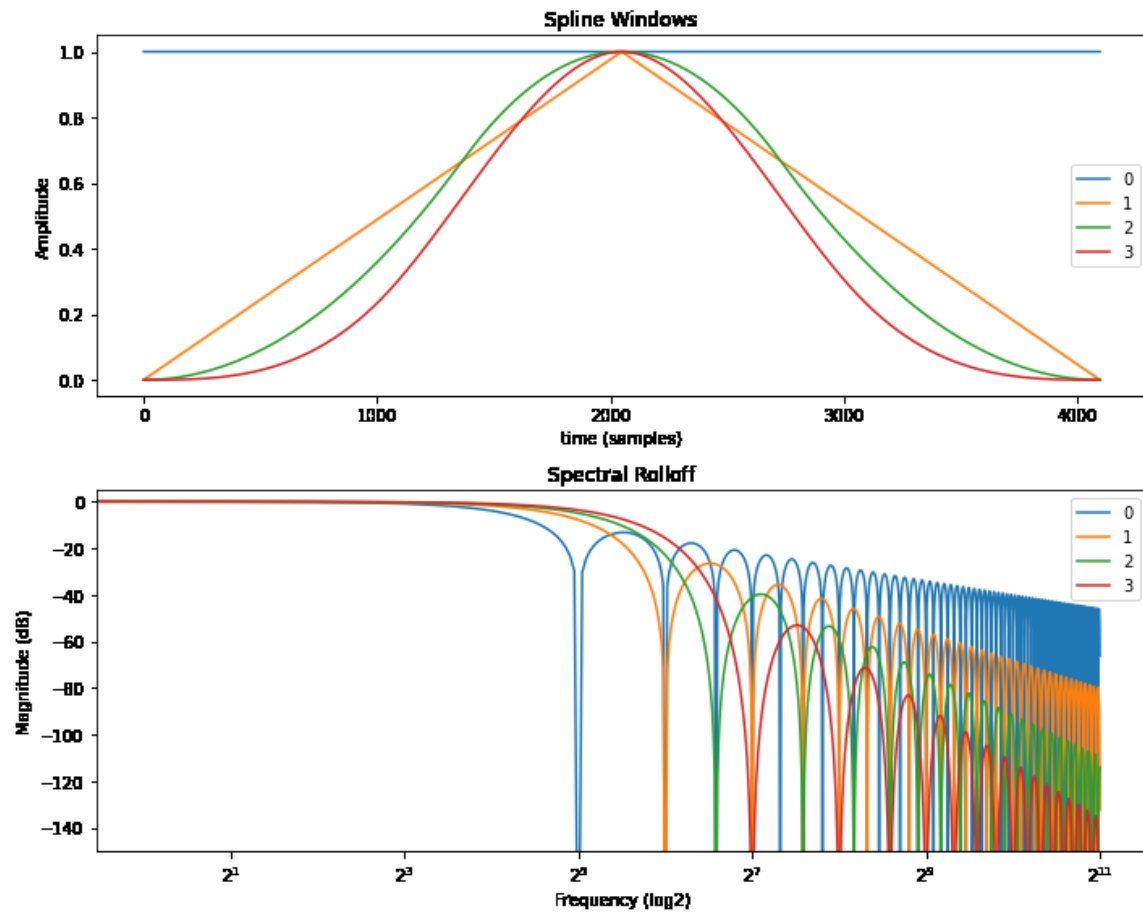
- $N = 0 \Rightarrow$  Rectangular (constant)
- $N = 1 \Rightarrow$  Triangular (linear)
- $N = 2 \Rightarrow$  Quadratic
- $N = 3 \Rightarrow$  Cubic

## Roll-Off Rate:

As  $N$  increases, the window becomes smoother.

$w_{\text{Spline}(N)}$  is  $(N - 1)$ -times continuously differentiable, and has roll-off rate  $6(N + 1)$  dB per octave.

# Spline Window Examples



# Miscellaneous Windows

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## Bartlett ( “Triangular” )

$$w(n) = w_R(n) \left[ 1 - \frac{|n|}{(M-1)/2} \right]$$

- Convolution of two half-length rectangular windows
- Window transform is  $\text{sinc}^2 \implies$ 
  - First sidelobe twice as far down as rect (-26 dB)
  - Main lobe twice as wide as that of a rectangular window *having the same length*  
(same as that of a half-length rect used to make it)
- Often applied to *sample correlations* of finite data
- Also called the “tent function”
- $M-1$  often replaced by  $M$  or  $M+1$  to avoid including endpoint zeros

## Using Any Window as a Tapering Function

Sometimes we need a wide rectangular window with *tapered edges*:

1. *Split* any window into halves, inserting the rectangle between
2. *Convolve* the rectangular window with any desired window

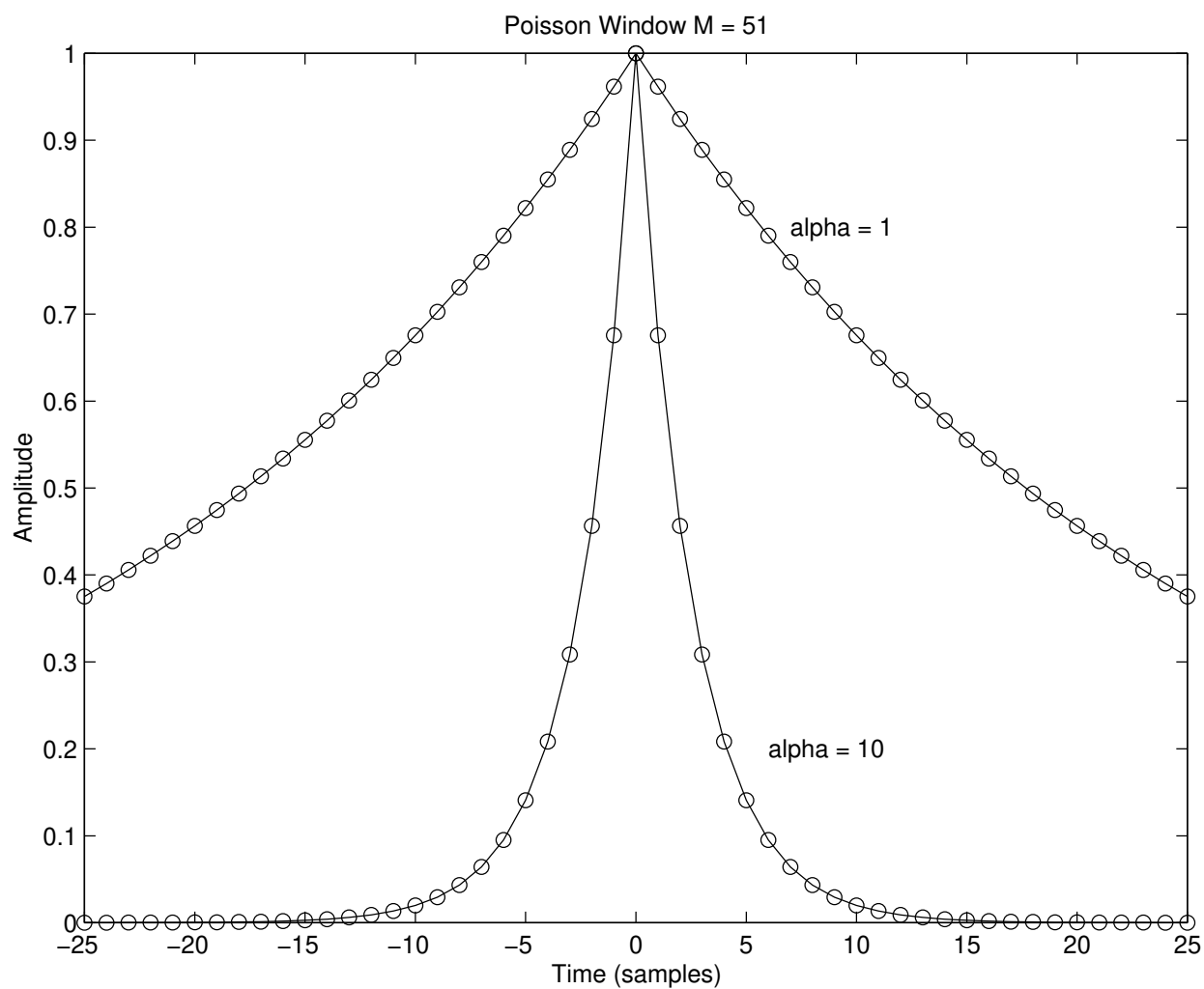
Method 2 preserves *smoothness* of the tapering window and hence its *roll-off rate*.

## Poisson (“Exponential”)

$$w_P(n) = w_R(n)e^{-\alpha \frac{|n|}{(M-1)/2}}$$

where  $\alpha$  determines the time constant  $\tau$ :

$$\frac{\tau}{T} = \frac{M-1}{2\alpha} \quad \text{samples}$$



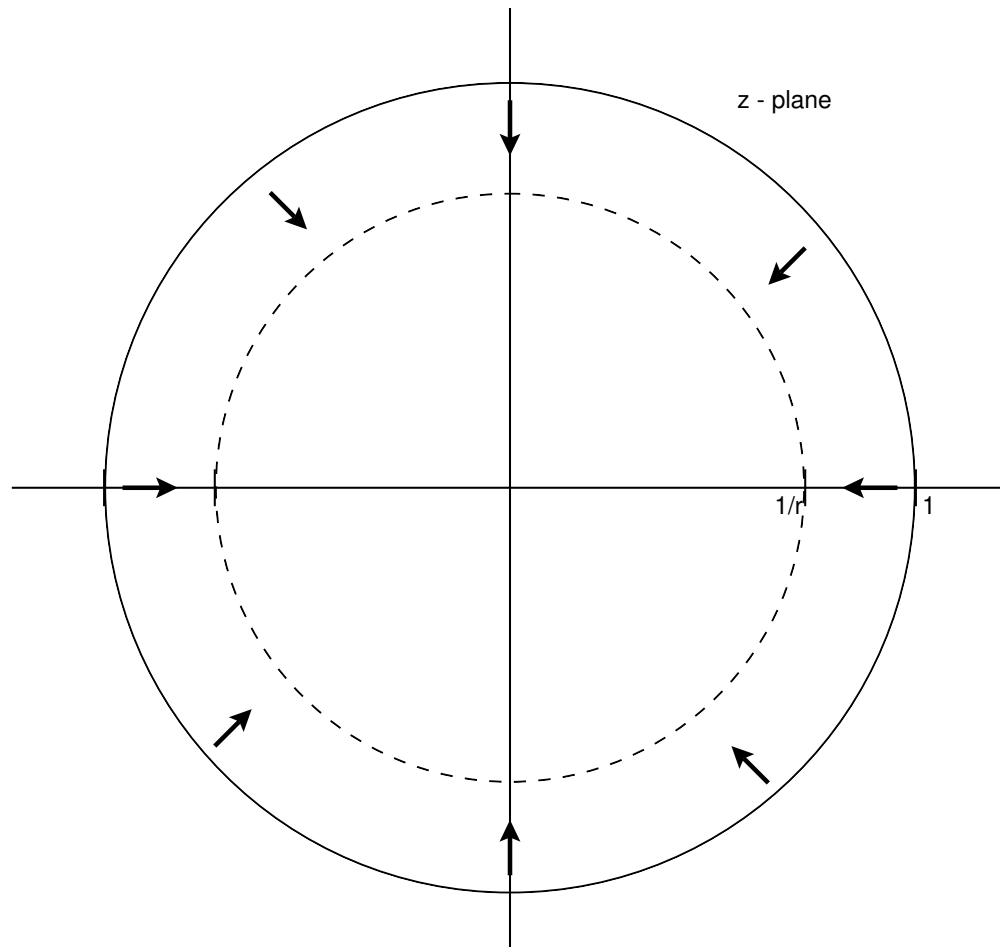
## Poisson Window in System Identification

In the  $z$ -plane, the Poisson window has the effect of contracting the spectrum toward zero inside unit circle. Consider an *infinitely long* Poisson window (no truncation by a rectangular window  $w_R$ ) applied to a causal signal  $h(n)$  having  $z$  transform  $H(z)$ :

$$\begin{aligned} H_P(z) &= \sum_{n=0}^{\infty} [w(n)h(n)]z^{-n} \\ &= \sum_{n=0}^{\infty} \left[ h(n)e^{-\frac{\alpha n}{M/2}} \right] z^{-n} \quad (\text{let } r \triangleq e^{\frac{\alpha}{M/2}}) \\ &= \sum_{n=0}^{\infty} h(n)z^{-n}r^{-n} = \sum_{n=0}^{\infty} h(n)(zr)^{-n} \\ &= H(zr) \end{aligned}$$

- Unit-circle response moved to  $|z| = 1/r < 1$
- Marginally stable poles now decay as  $r^{-n} = e^{-\alpha n/(M/2)}$





The Poisson window can be useful for impulse-response modeling by poles and/or zeros ("system identification"). In such applications, the window length is best chosen to include substantially all of the impulse-response data.

## Hann-Poisson (“No Sidelobes”)

$$w(n) = \frac{1}{2} \left[ 1 + \cos \left( \pi \frac{n}{(M-1)/2} \right) \right] e^{-\alpha \frac{|n|}{(M-1)/2}}$$

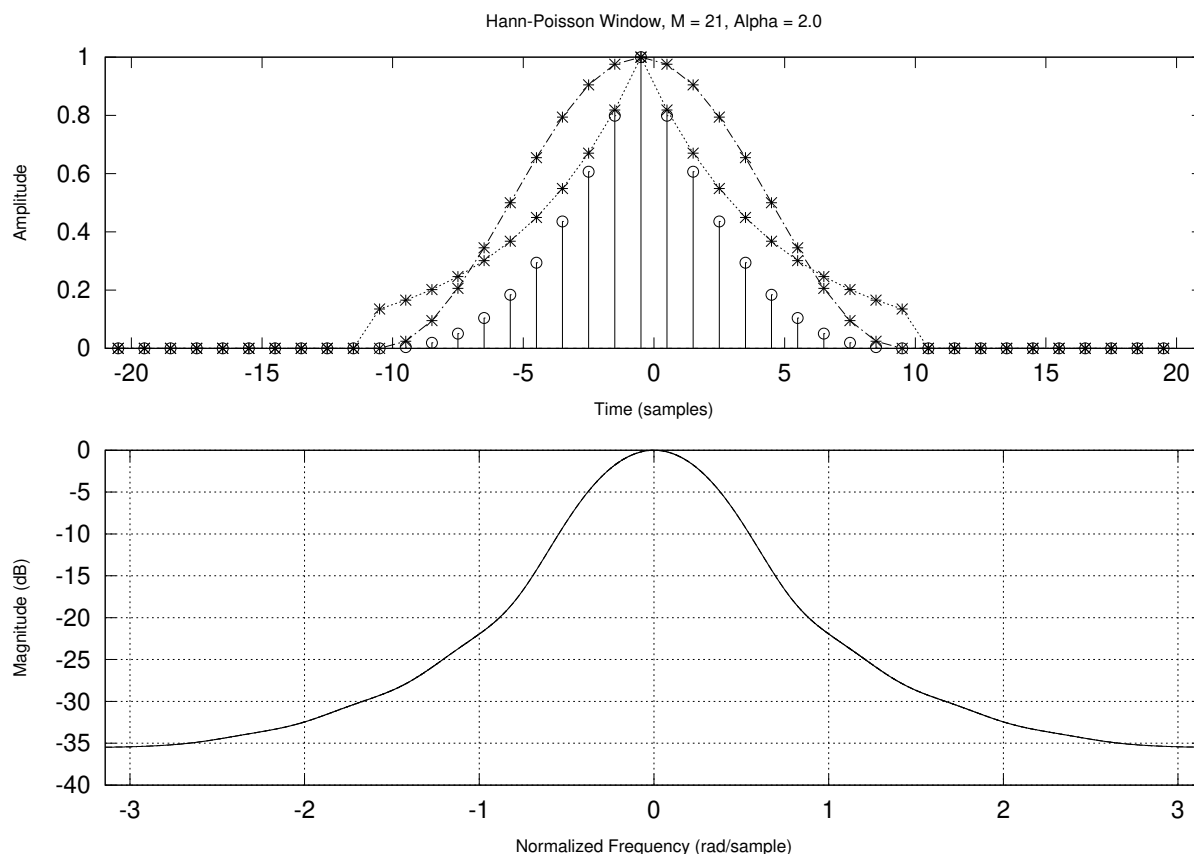
- Poisson window times Hann window (exponential times raised cosine)
- “No sidelobes” for  $\alpha \geq 2$
- Valuable for “hill climbing” optimization methods (gradient-based)

Matlab:

```
function [w,h,p] = hannpoisson(M,alpha)
%HANNPOISSON - Length M Hann-Poisson window

Mo2 = (M-1)/2; n=(-Mo2:Mo2)';
scl = alpha / Mo2;
p = exp(-scl*abs(n));
scl2 = pi / Mo2;
h = 0.5*(1+cos(scl2*n));
w = p.*h;
```

# Hann-Poisson Window and Transform



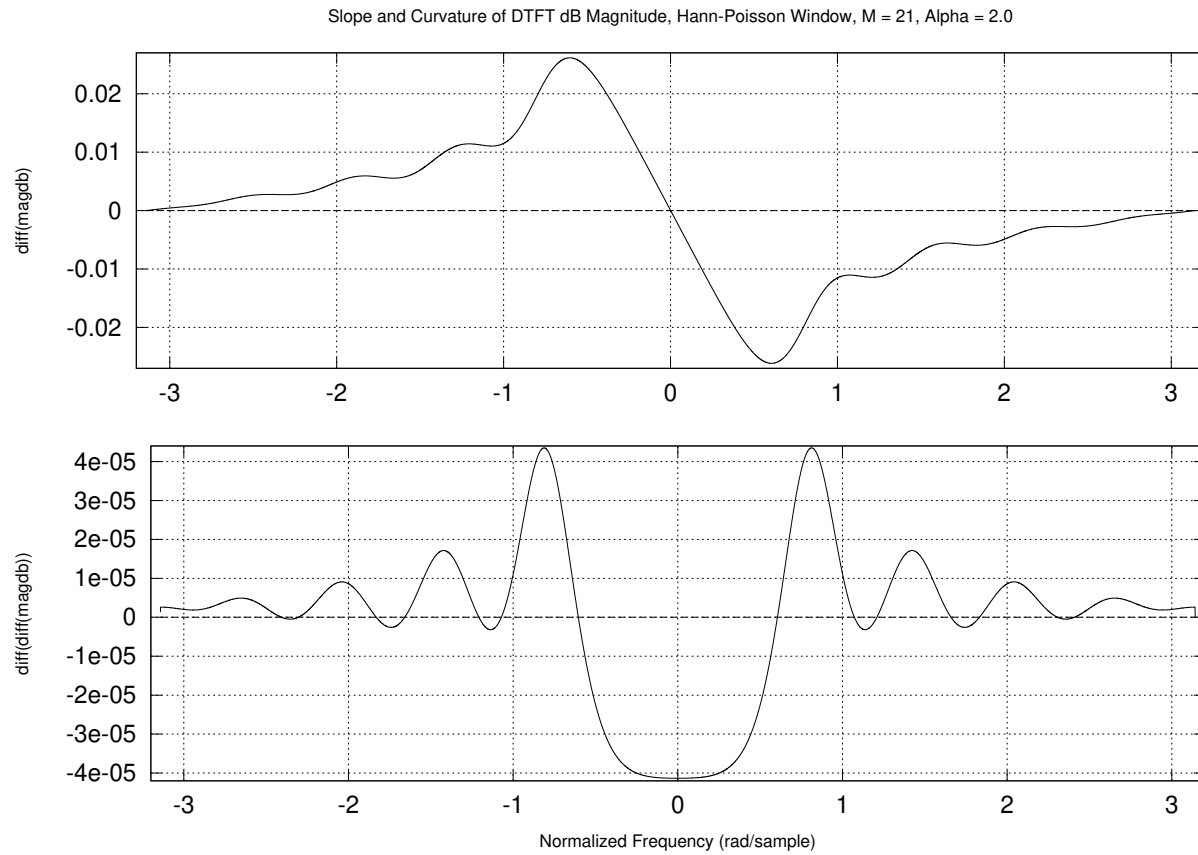
[https://ccrma.stanford.edu/~jos/Windows/Hann\\_Poisson\\_Window\\_Transform.html](https://ccrma.stanford.edu/~jos/Windows/Hann_Poisson_Window_Transform.html)

## Question:

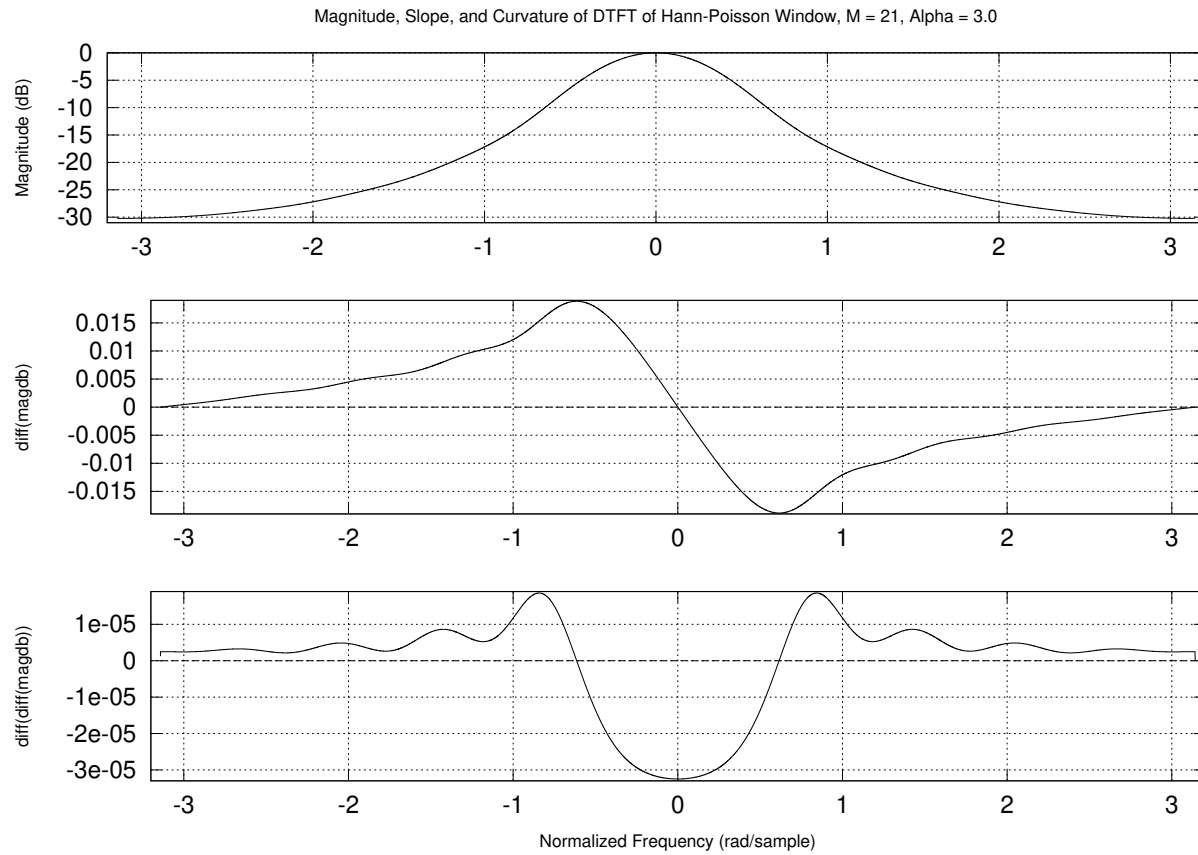
How can a truncated *anything* avoid having ripples in its spectrum? (“Gibbs ripples”)

Let’s look at the *derivatives* of the window:

# Hann-Poisson Slope and Curvature



# Slope and Curvature for Larger Alpha



## Maximum Main-Lobe Energy Window: DPSS

**Question:** How do we use *all*  $M$  degrees of freedom (sample values) in an  $M$ -point window  $w(n)$  to obtain  $W(\omega) \approx \delta(\omega)$  in some optimal sense?

That is, we wish to perform the following optimization:

$$\max_w \left[ \frac{\text{main lobe energy}}{\text{total energy}} \right]$$

In the continuous-time case  $[\omega \in (-\infty, \infty)]$ , this problem is solved by a *prolate spheroidal wave function*, an eigenfunction of the integral equation

$$\int_{-\omega_c}^{\omega_c} W(\nu) \frac{\sin[\pi D \cdot (\omega - \nu)]}{\pi(\omega - \nu)} d\nu = \lambda W(\omega), \quad |\omega| \leq \omega_c$$

where  $D$  is the nonzero time-duration of  $w(t)$  in seconds.

**Interpretation:**

$$\begin{aligned} & [\text{CHOP}_{2\omega_c}(W)] * [D \text{sinc}(D\omega)] \\ &= \text{FT}(\text{CHOP}_D(\text{IFT}(\text{CHOP}_{2\omega_c}(W)))) = \lambda W \end{aligned}$$

where  $\text{CHOP}_D(w)$  is a rectangular windowing operation which zeros  $w$  outside the interval  $t \in [-D/2, D/2]$ .

$W$  is thus the *bandlimited extrapolation* of its main lobe ( $\omega \in [-\omega_c, \omega_c]$ )

The optimal window transform  $W$  is an eigenfunction of this operation sequence corresponding to the *largest eigenvalue*.

The resulting optimal window  $w$  has maximum main-lobe energy as a fraction of total energy.

It may be called the *Slepian window*, or *prolate spheroidal window* in the continuous-time case.

In discrete time, we need *Discrete Prolate Spheroidal Sequences (DPSS)*, eigenvectors of the following symmetric Toeplitz matrix constructed from a sampled sinc function:

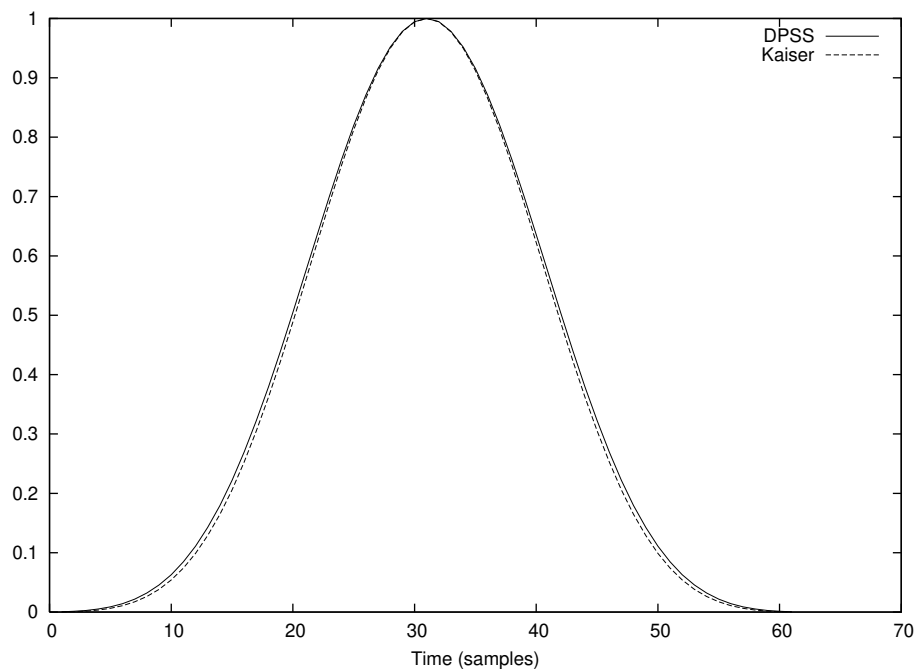
$$S[k, l] = \frac{\sin[\omega_c T(k - l)]}{k - l}, \quad k, l = 0, 1, 2, \dots, M - 1$$

- $M$  = window length in samples
- $\omega_c$  = main-lobe cut-off frequency (rad/sec)
- $T$  = sampling period in seconds.

The DPSS window (digital Slepian window) is then given by the eigenvector corresponding to the largest eigenvalue.

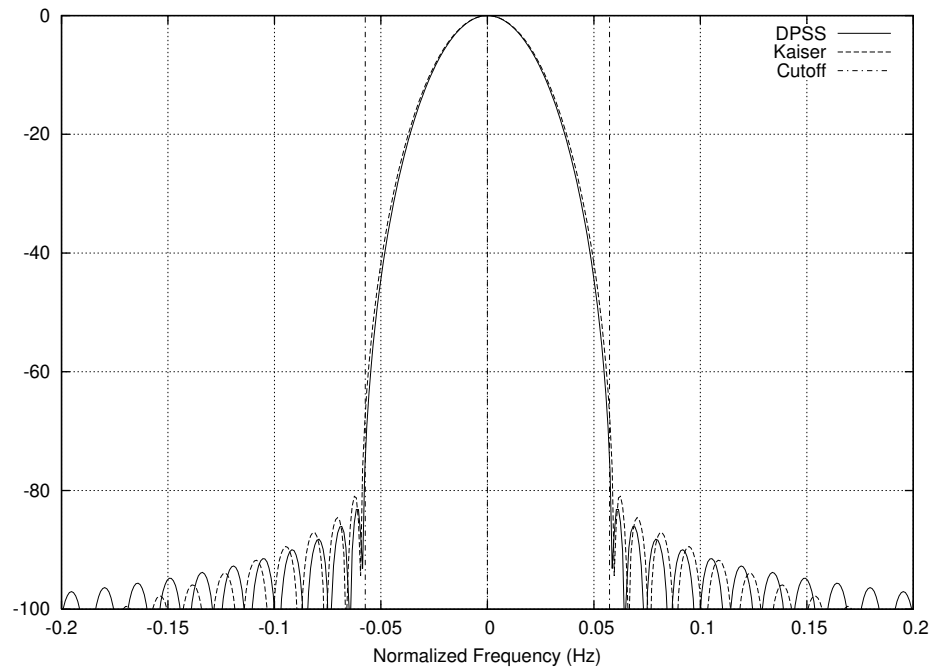
## Matlab for the DPSS Window

```
function [w,A,V] = dpssw(M,Wc);  
% DPSSW - Digital Prolate Spheroidal Sequence window  
%          (Slepian window) of length M, having  
%          cut-off frequency Wc in (0,pi).  
k = (1:M-1);  
s = sin(Wc*k)./ k;  
c0 = [Wc,s];  
A = toeplitz(c0); % c0=1st col of symm. Toeplitz  
[V,evals] = eig(A); % Only need principal eigenvector  
[emax,imax] = max(abs(diag(evals)));  
w = V(:,imax);  
w = w / max(w);
```





# Kaiser and DPSS Window Transforms Compared



- Kaiser  $\approx$  DPSS
- DPSS window has a slightly narrower main lobe
- DPSS window has lower overall side-lobe levels
- Kaiser window side lobes roll off faster
- Otherwise they are very similar

## Kaiser (Kaiser-Bessel)

Kaiser discovered a very good approximation to prolate spheroidal wave functions using Bessel functions:

$$w_K(n) \triangleq \begin{cases} \frac{I_0\left(\beta\sqrt{1-\left(\frac{n}{M/2}\right)^2}\right)}{I_0(\beta)}, & -\frac{M-1}{2} \leq n \leq \frac{M-1}{2} \\ 0, & \text{elsewhere} \end{cases}$$
$$w_K(n) \triangleq w_R(n) \frac{I_0\left(\beta\sqrt{1-\left(\frac{n}{M/2}\right)^2}\right)}{I_0(\beta)}$$

This is called the *Kaiser* (or *Kaiser-Bessel*) window.

The Fourier transform of the Kaiser window  $w_K(t)$  (where  $t$  is treated as continuous) is given by

$$\begin{aligned} W(\omega) &= \frac{M}{I_0(\beta)} \frac{\sinh\left(\sqrt{\beta^2 - \left(\frac{M\omega}{2}\right)^2}\right)}{\sqrt{\beta^2 - \left(\frac{M\omega}{2}\right)^2}} \\ &= \frac{M}{I_0(\beta)} \frac{\sin\left(\sqrt{\left(\frac{M\omega}{2}\right)^2 - \beta^2}\right)}{\sqrt{\left(\frac{M\omega}{2}\right)^2 - \beta^2}} \end{aligned}$$

where  $I_0$  is the zero-order modified Bessel function of the first kind.

## Modified Bessel Function of the 1st Kind

A *series expansion* for the order zero, modified Bessel function of the first kind is given by

$$I_0(x) = \sum_{k=0}^{\infty} \left[ \frac{(x/2)^k}{k!} \right]^2$$

Compare this with

$$e^{x/2} = \sum_{k=0}^{\infty} \frac{(x/2)^k}{k!}.$$

## Kaiser-Bessel Window Notes

$$w_K(n) \triangleq w_R(n) \frac{I_0 \left( \beta \sqrt{1 - \left( \frac{n}{M/2} \right)^2} \right)}{I_0(\beta)}$$

- “Closed form” (given  $I_0$  series or table)
- Reduces to rectangular window for  $\beta = 0$
- Asymptotic roll-off is 6 dB/octave
- For  $\beta \gg 0$ , first null in window transform is at  
 $\omega_0 \approx 2\beta/M$   
 $\Rightarrow \beta = M\omega_0/2$
- Sometimes the Kaiser window is parameterized by  $\alpha$ :

$$\beta \triangleq \pi\alpha$$

## Kaiser Window Time-Bandwidth Product

- Define the main-lobe “cutoff frequency” as half-way to the first zero in  $W(\omega)$ :

$$\omega_c = \frac{\omega_0}{2} = \frac{\beta}{M} = \frac{\pi\alpha}{M}$$

- Then

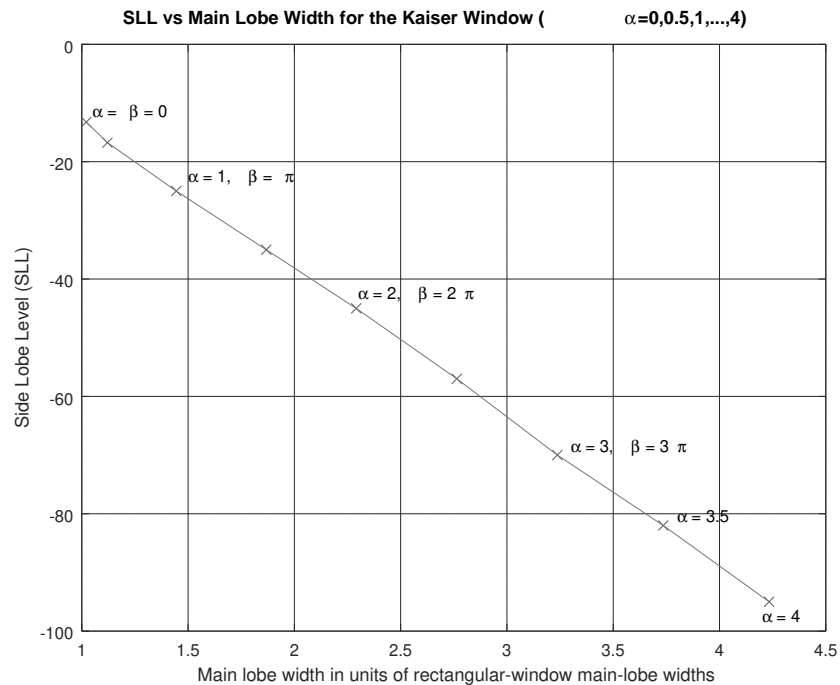
$$\begin{aligned}\beta &= M\omega_c = \frac{1}{2}M \cdot (2\omega_c) \\ &= \frac{1}{2} \text{ duration (samples)} \times \text{bandwidth (rad/sample)} \\ \alpha &= \frac{\beta}{\pi} = \frac{2\beta}{2\pi} \\ &= \text{duration (samples)} \times \text{bandwidth (cycles/sample)}\end{aligned}$$

- $\beta = M\omega_c T$  is equal to 1/2 ‘time-bandwidth product’  
 $\beta = \frac{1}{2}\Delta t \cdot \Delta\omega \Rightarrow \alpha = \Delta t \cdot \Delta f$

- In this definition of time-bandwidth product, the “cut-off frequency  $\omega_c$  of the Kaiser-window transform is defined as *half* of the first null frequency, *i.e.*,  
 $\omega_c = \omega_0/2$ .

- $\beta$  trades off side lobe level for main lobe width  
larger  $\beta \Rightarrow$  lower S.L.L., wider mainlobe

# Kaiser Side-Lobe Level vs. Main-Lobe Width

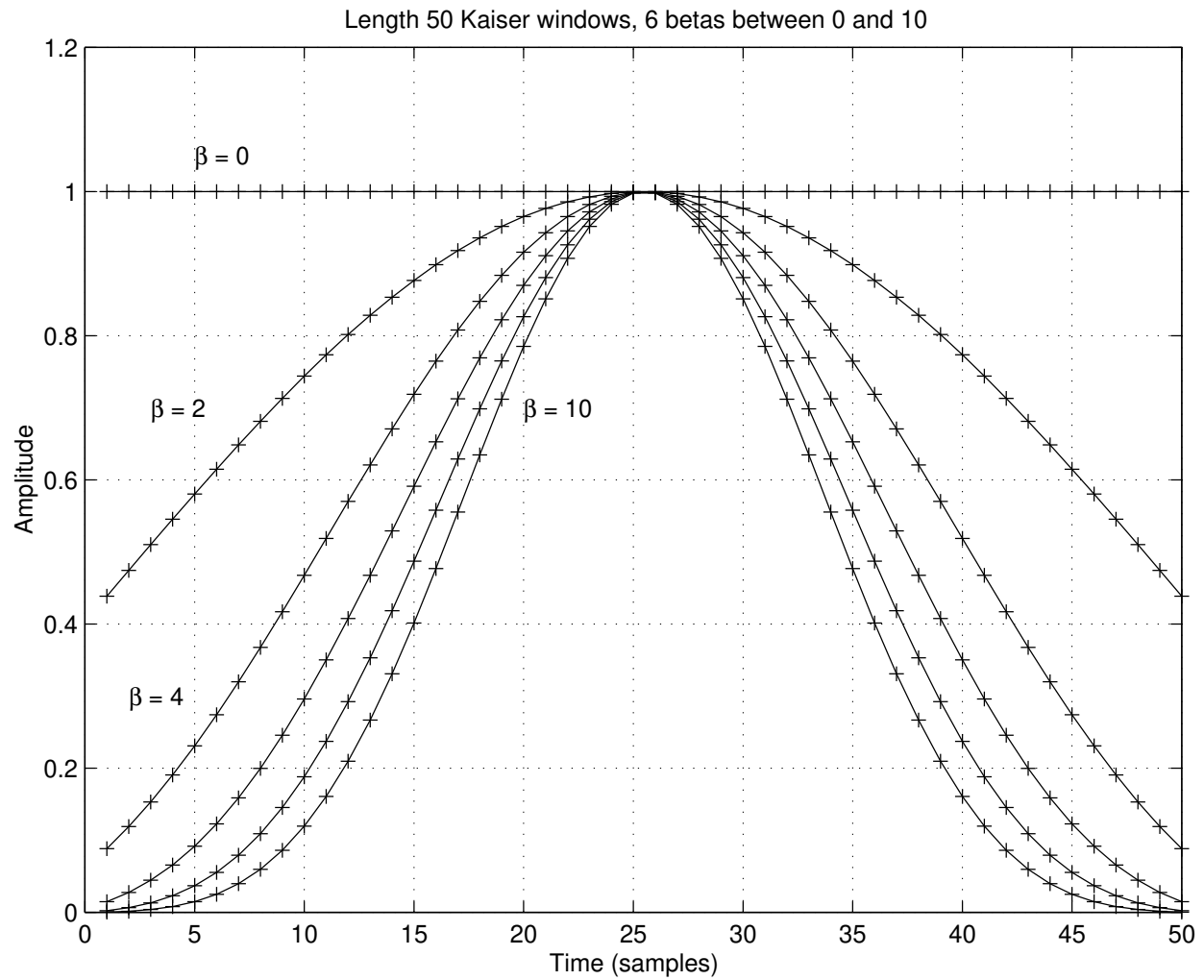


## Mathematica Demonstration:

<http://demonstrations.wolfram.com/KaiserWindowTransform/>

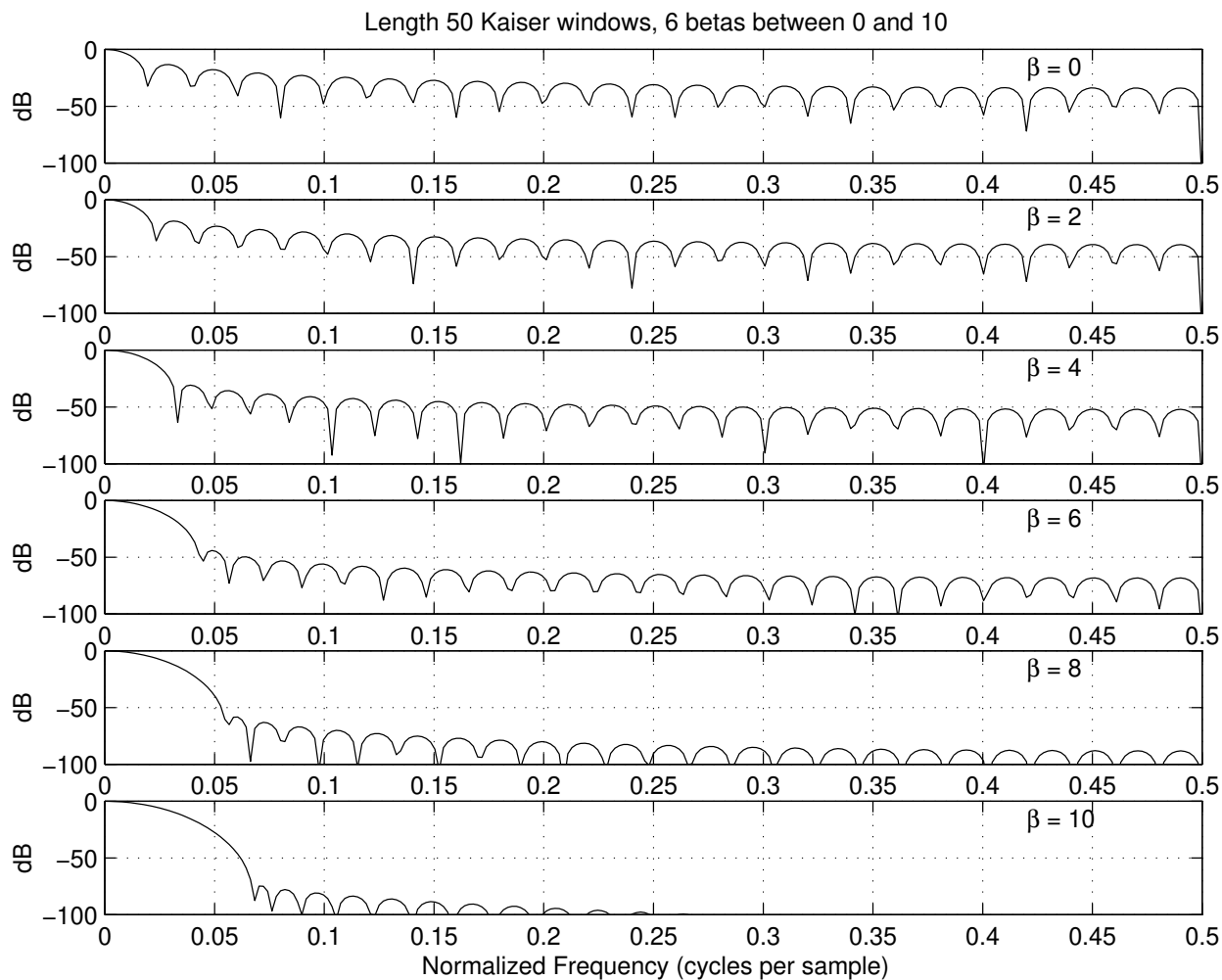
# Kaiser Window Examples

$$\beta = [0, 2, 4, 6, 8, 10]$$



# Kaiser Window Transform Examples

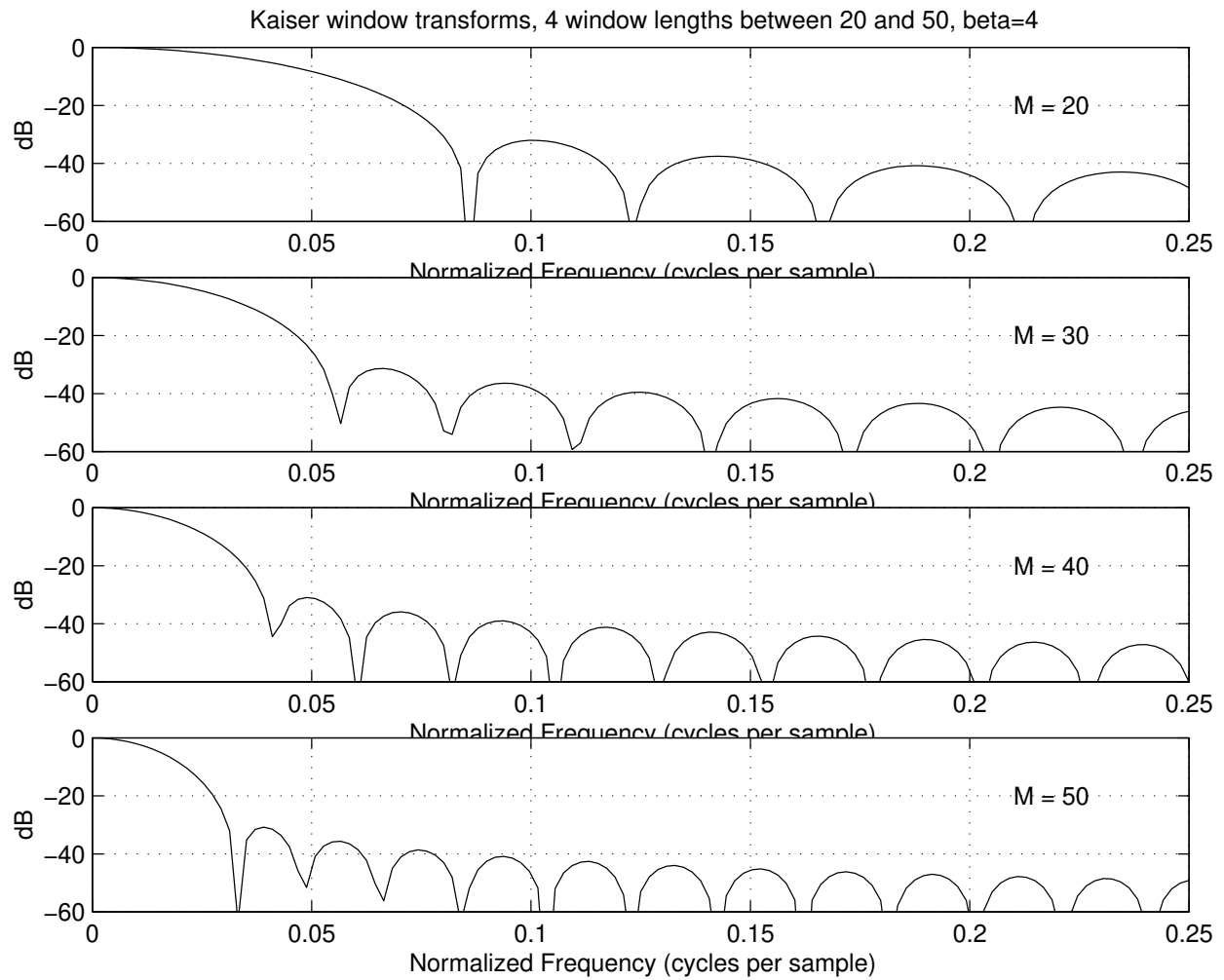
$$\beta = [0, 2, 4, 6, 8, 10]$$





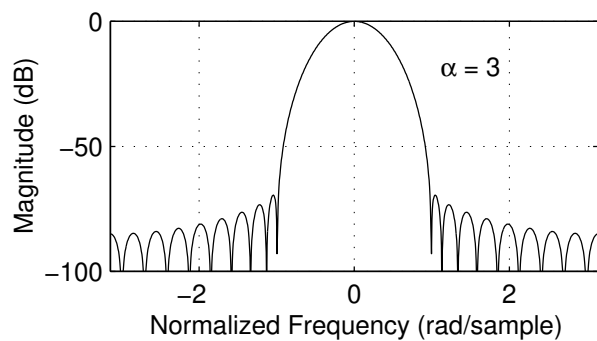
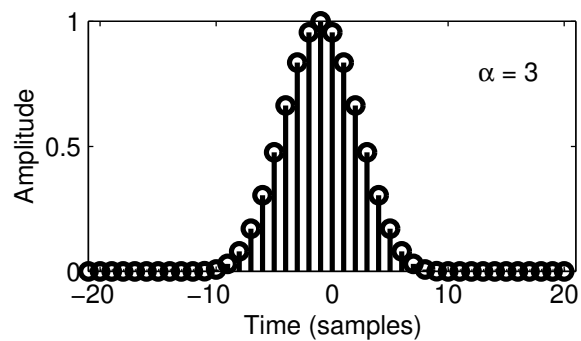
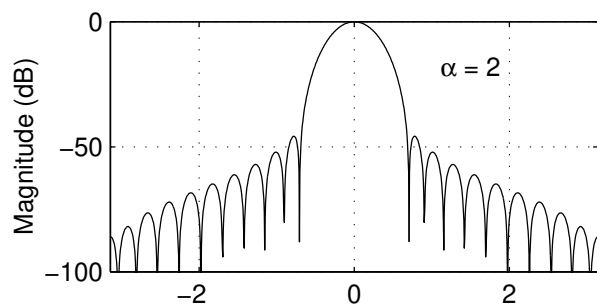
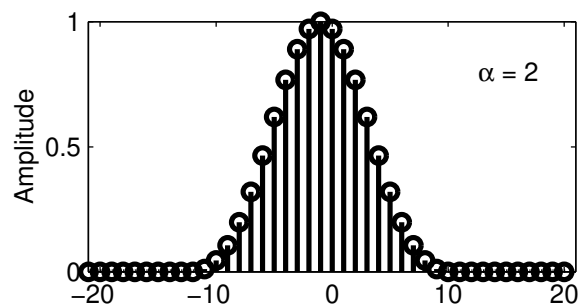
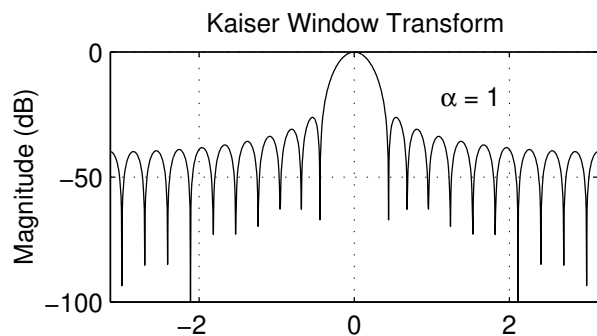
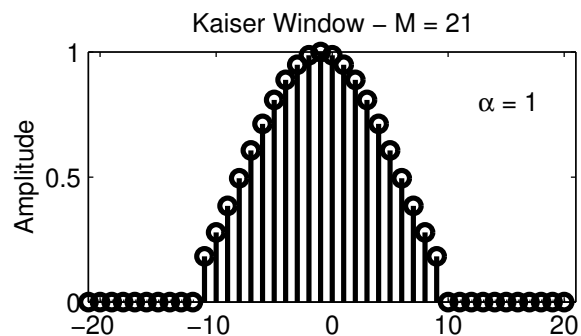
# Kaiser Window: Different Lengths M

$$M = [20, 30, 40, 50]$$

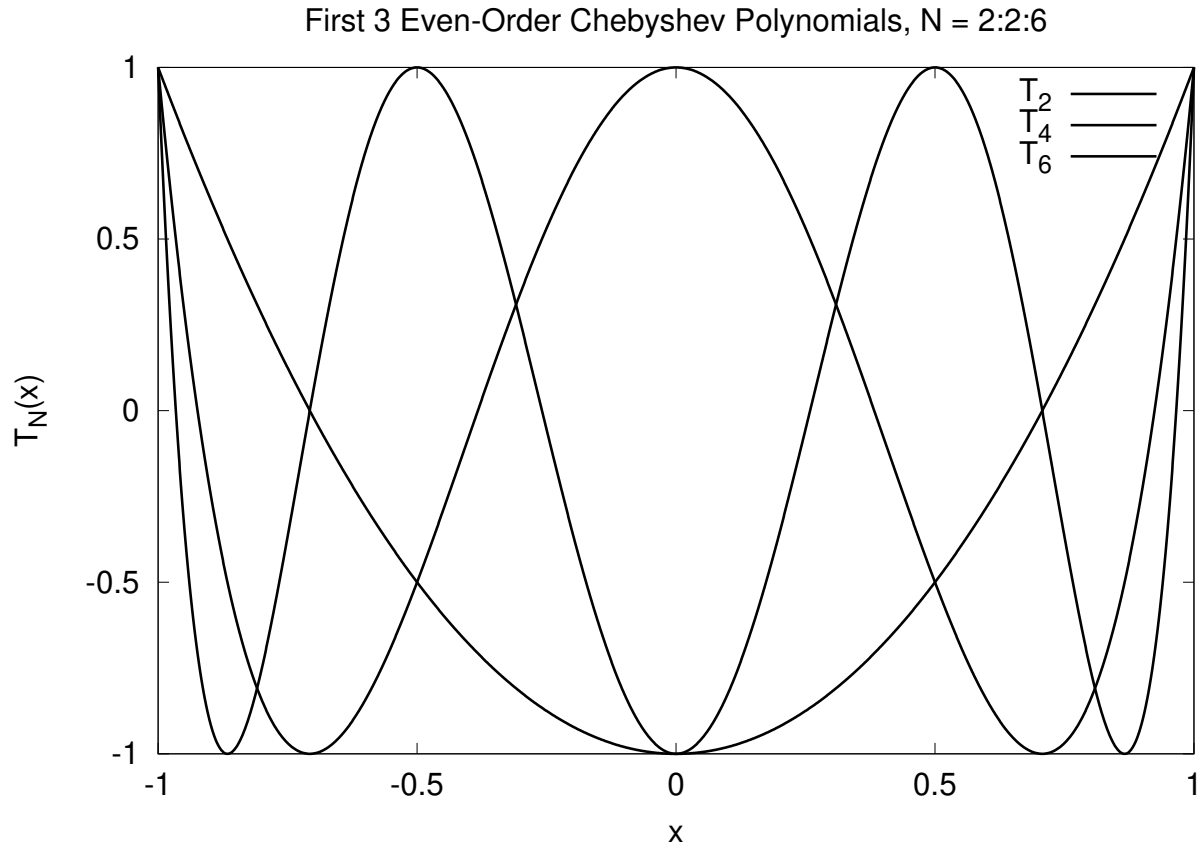


# More Kaiser-Window Examples

$$\alpha = [1, 2, 3] \quad (\beta = [\pi, 2\pi, 3\pi])$$



# Chebyshev Polynomials



The  $n$ th *Chebyshev polynomial* may be defined by

$$T_n(x) = \begin{cases} \cos[n \cos^{-1}(x)], & |x| \leq 1 \\ \cosh[n \cosh^{-1}(x)], & |x| > 1 \end{cases}.$$

Clearly,  $T_0(x) = 1$  and  $T_1(x) = x$ .

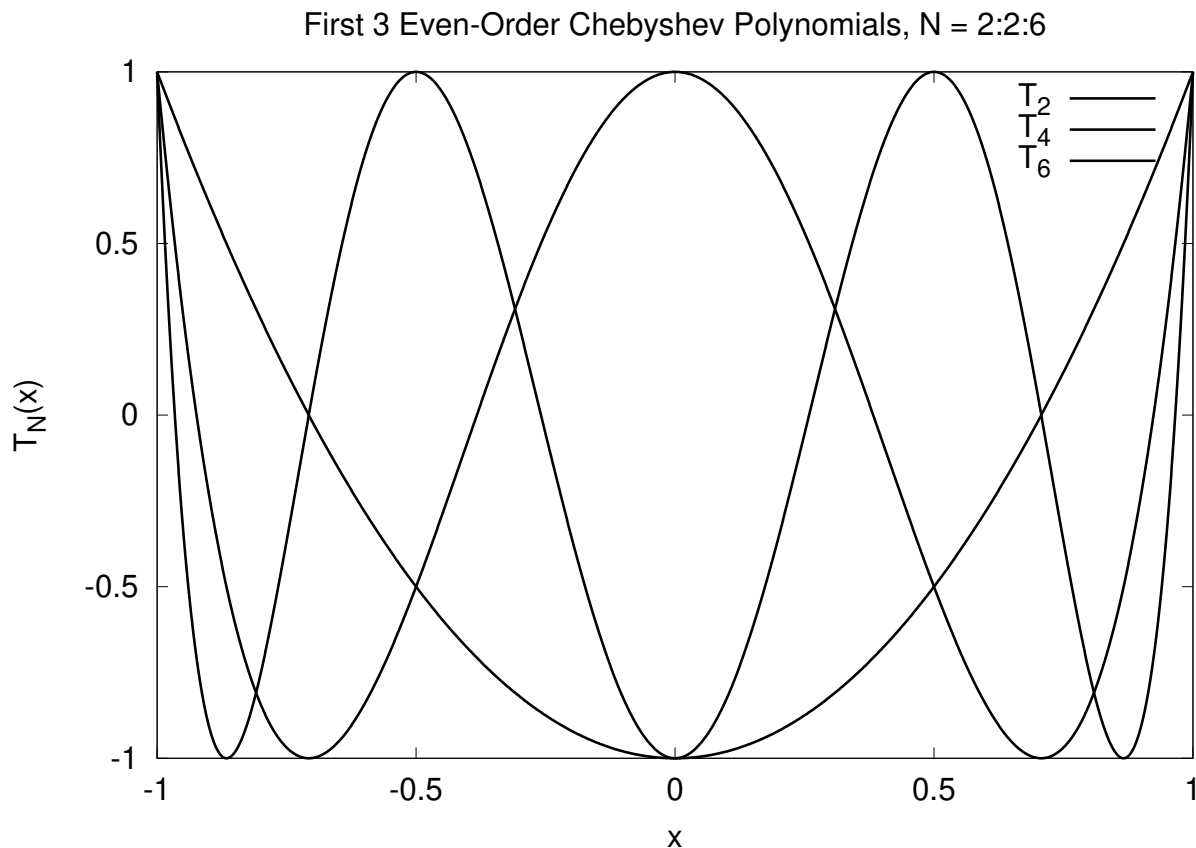
Using the double-angle trig formula

$\cos(2\theta) = 2 \cos^2(\theta) - 1$ , it can be verified that

$$T_n(x) = 2xT_{n-1}(x) - T_{n-2}(x) \quad (n \geq 2)$$

# Chebyshev Polynomial Properties

- $T_n(x)$  is an  $n$ th-order polynomial in  $x$
- $T_n(x)$  is an even function when  $n$  is an even integer, and odd when  $n$  is odd
- $T_n(x)$  has  $n$  zeros in the open interval  $(-1, 1)$ , and  $n + 1$  extrema in the closed interval  $[-1, 1]$
- $T_n(x) > 1$  for  $x > 1$



$$T_n(x) = \cos[n \cos^{-1}(x)], \quad |x| \leq 1$$

## Dolph-Chebyshev Window

Minimize the *Chebyshev norm* of the side lobes, e.g.,

$$\begin{aligned} & \min_{w, \sum w=1} \| \text{sidelobes}(W) \|_{\infty} \\ & \equiv \min_{w, \sum w=1} \{ \max_{\omega > \omega_c} |W(\omega)| \} \end{aligned}$$

Alternatively, *minimize main lobe width* subject to a sidelobe spec:

$$\min_{w, W(0)=1} (\omega_c) \quad \left| \quad |W(\omega)| \leq c_{\alpha}, \forall |\omega| \geq \omega_c \right.$$

Closed-Form Window Transform (Dolph):

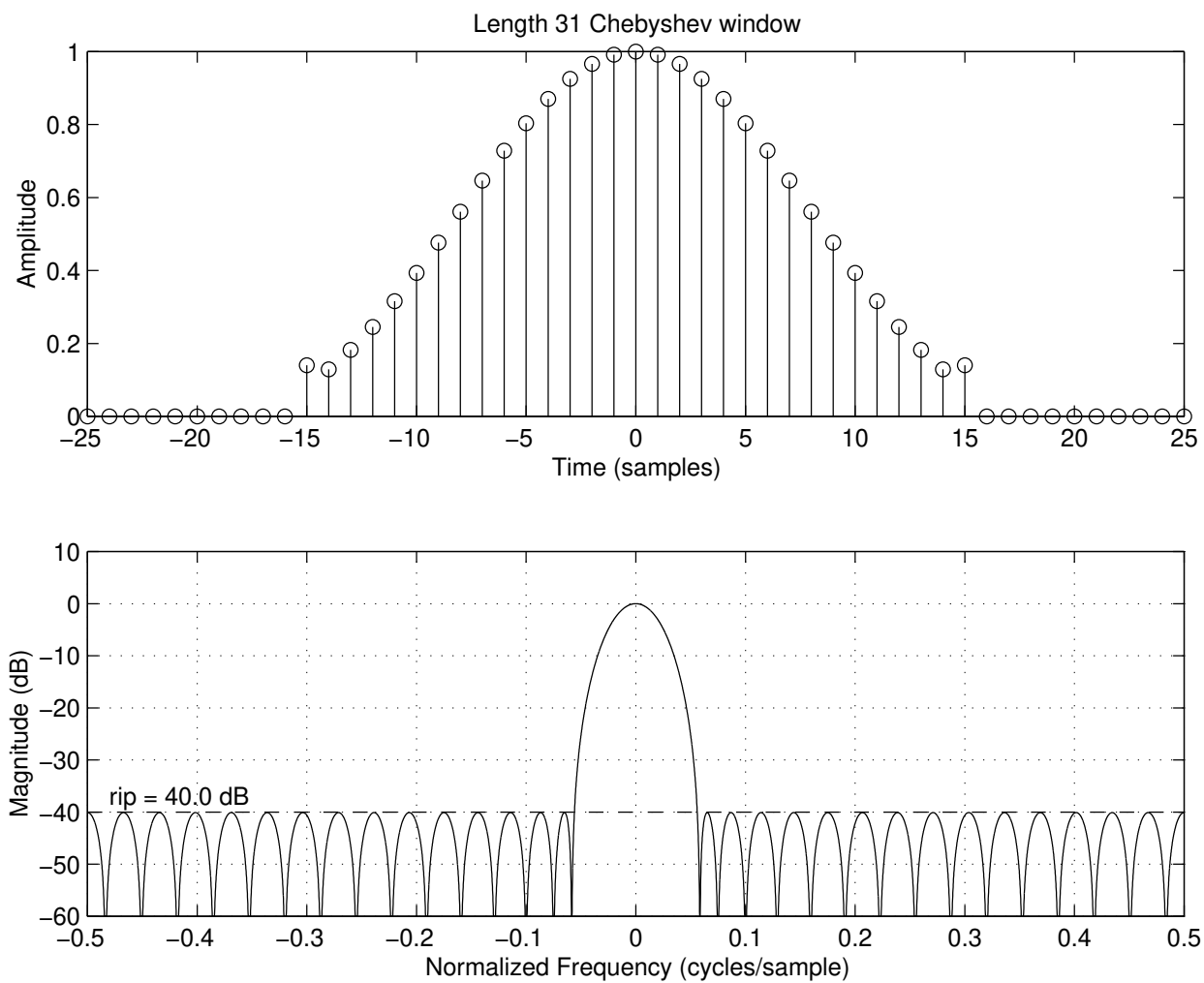
$$\begin{aligned} W(\omega_k) &= \frac{\cos \left\{ M \cos^{-1} \left[ \Gamma \cos \left( \frac{\pi k}{M} \right) \right] \right\}}{\cosh \left[ M \cosh^{-1}(\Gamma) \right]}, \quad (|k| \leq M-1) \\ \Gamma &= \cosh \left[ \frac{1}{M} \cosh^{-1}(10^{\alpha}) \right] \geq 1, \quad (\alpha \approx 2, 3, 4) \end{aligned}$$

- Window  $w = \text{IDFT}(W)$  [zero-centered case]  
or IDFT of  $(-1)^k W(\omega_k)$  for causal case
- $\alpha$  controls sidelobe level (“stopband ripple”):

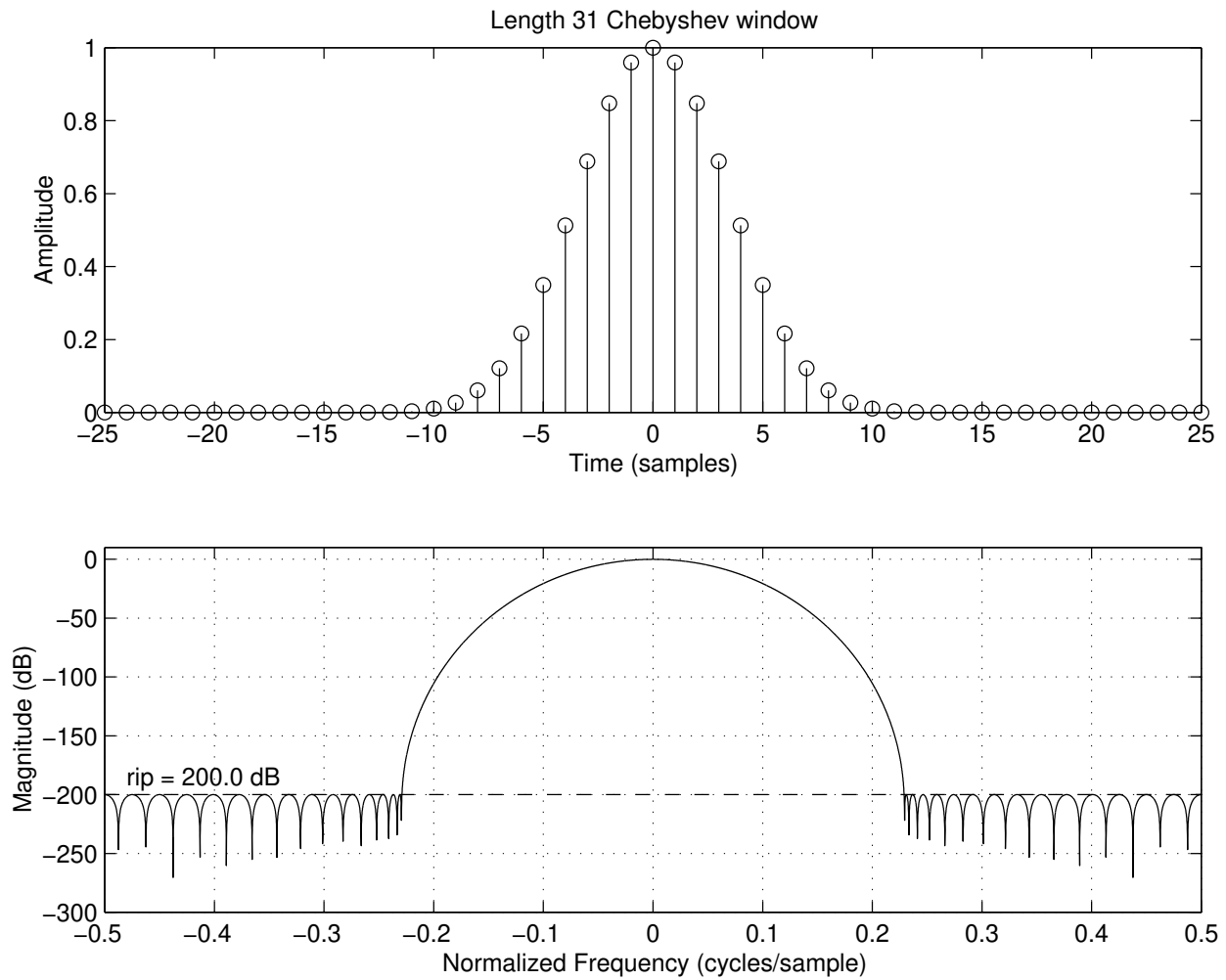
$$\text{Side-Lobe Level in dB} = -20\alpha.$$

- smaller ripple  $\Rightarrow$  larger  $\omega_c$
- see matlab function “chebwin(M,ripple)”

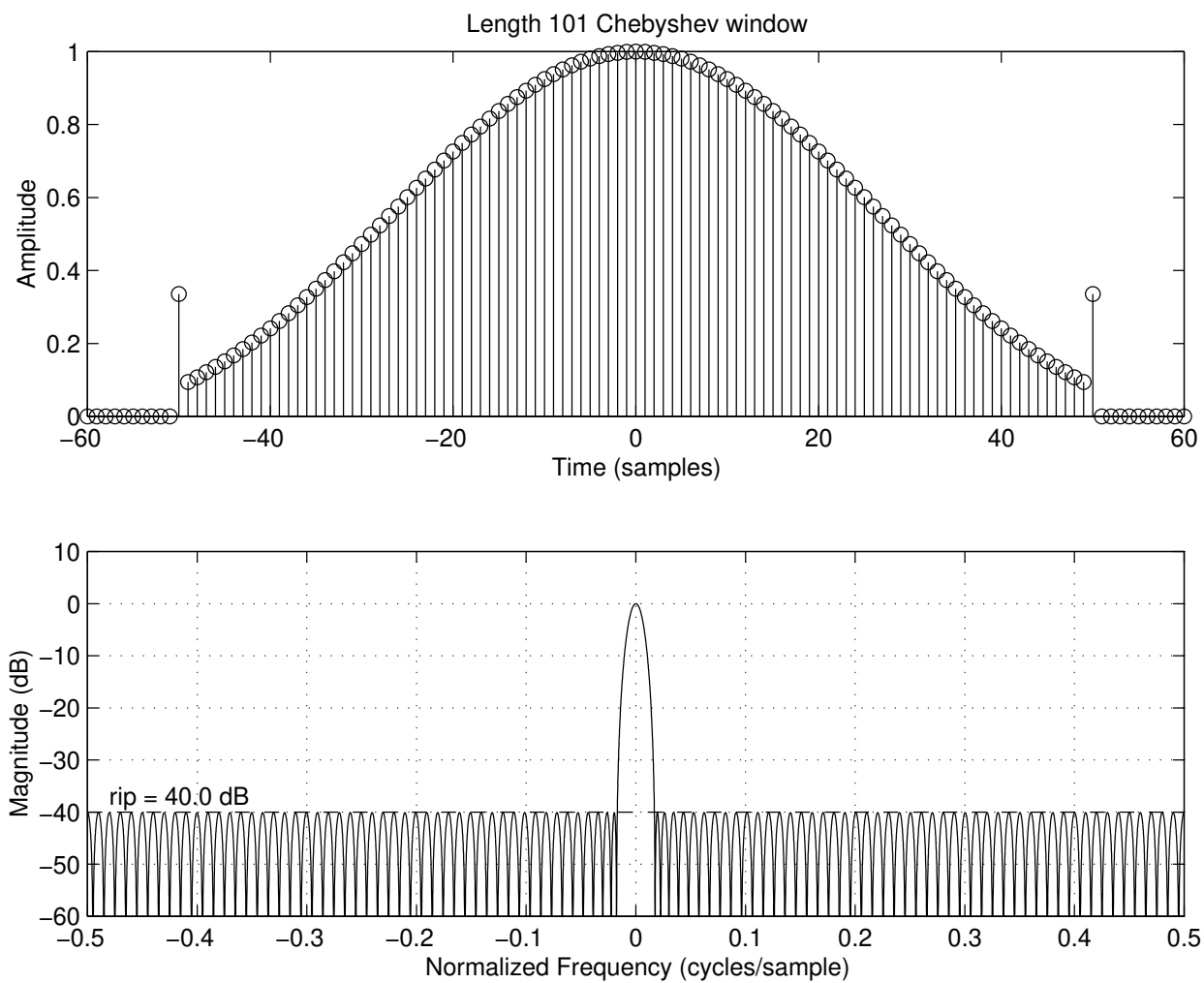
# Dolph-Chebyshev Window, Length 31, Ripple -40 dB



# Dolph-Chebyshev Window, Length 31, Ripple -200 dB

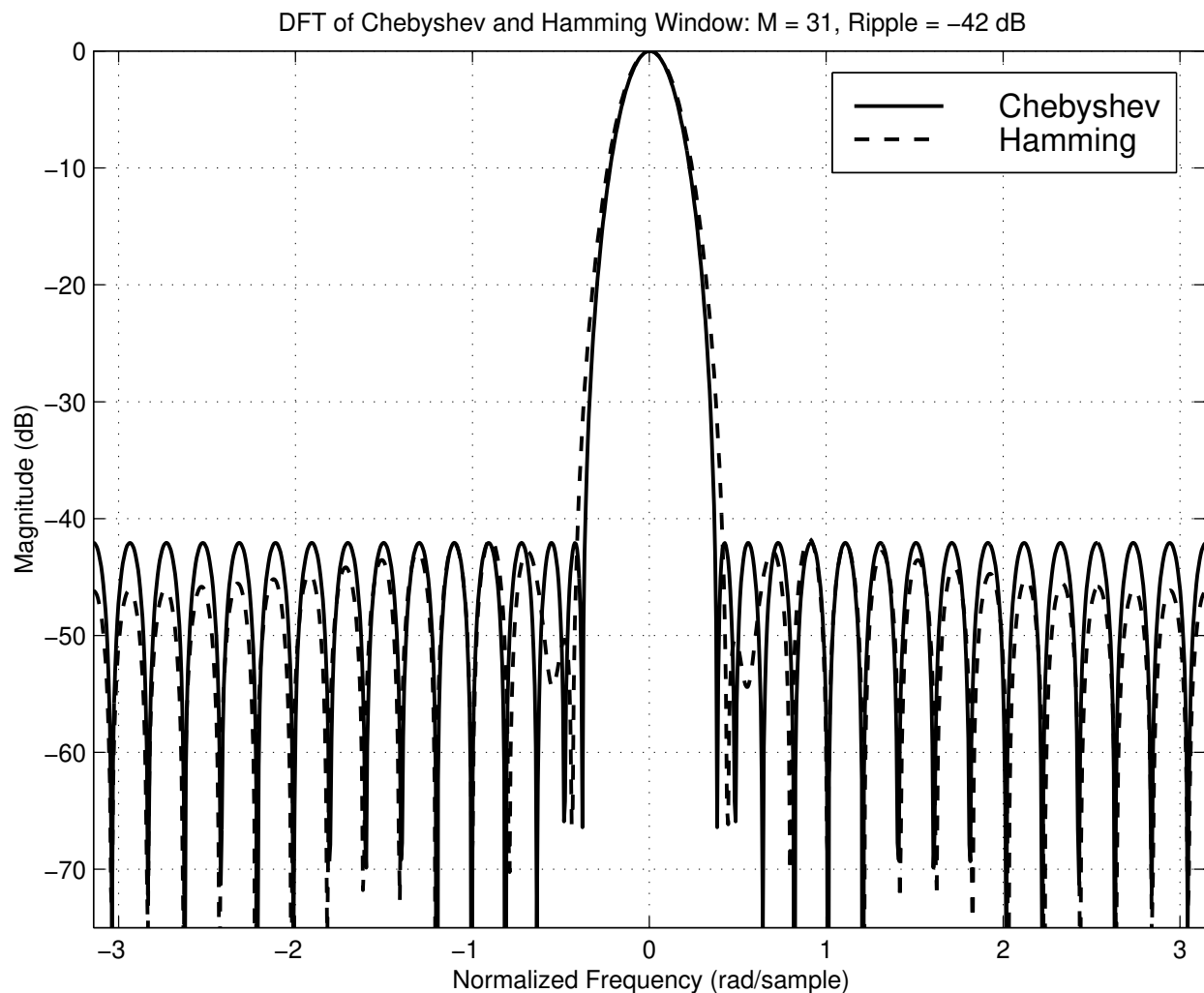


# Dolph-Chebyshev Window, Length 101, SLL -40 dB





# Dolph-Chebyshev and Hamming Windows Compared



For the comparison, we set the ripple parameter for `chebwin` to 42 dB:

```
window = [ chebwin(31,42)' zeros(1,1024-31) ];
```

## Gaussian

The Gaussian “bell curve” is the only smooth function that transforms to itself:

$$\frac{1}{\sigma\sqrt{2\pi}}e^{-t^2/2\sigma^2} \leftrightarrow e^{-\omega^2/2(1/\sigma)^2}$$

It also achieves the *minimum time-bandwidth product*

$$\sigma_t\sigma_\omega = \sigma \times (1/\sigma) = 1$$

when “width” of a function is defined as the square root of its second central moment. For even functions  $w(t)$ ,

$$\sigma_t \triangleq \sqrt{\int_{-\infty}^{\infty} t^2 w(t) dt}.$$

- Since the true Gaussian function has infinite duration, in practice we must *window it* with some finite window.
- Philippe Depalle suggests using a *triangular window* raised to some power  $\alpha$  for this purpose.
  - This choice *preserves the absence of sidelobes* for sufficiently large  $\alpha$ .
  - It also preserves *non-negativity* of the transform

## The Gaussian Window in Spectral Modeling

**Special Property:** On a dB scale, the Gaussian is *quadratic*  $\Rightarrow$  *parabolic interpolation of a sampled Gaussian transform is exact*.

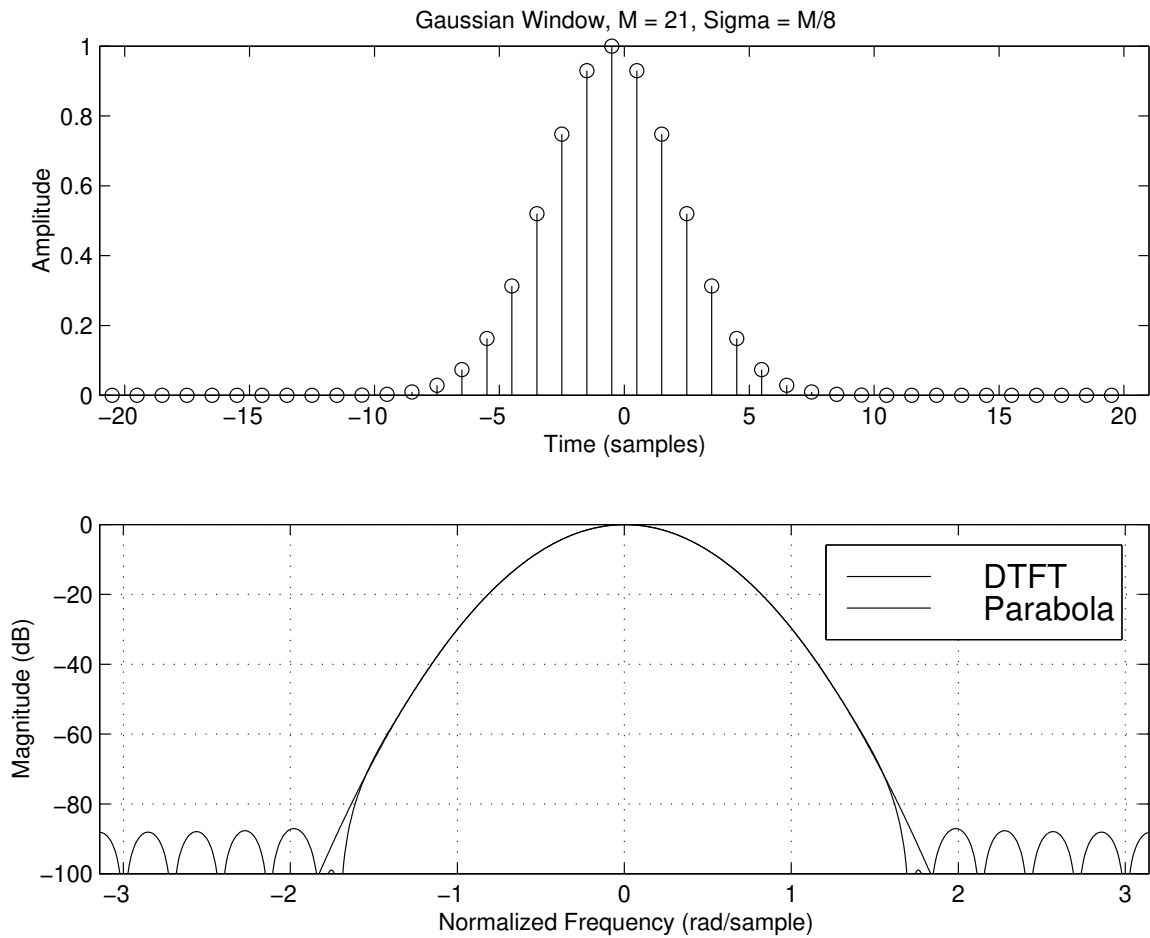
**Conjecture:** Quadratic interpolation of spectral peaks is generally more accurate on a *log-magnitude scale* (e.g., dB) than on a linear magnitude scale. This has been verified in a number of cases, and no counter-examples are yet known. **Exercise:** Prove this is true for the rectangular window.

### Matlab for the Gaussian Window

```
function [w] = gausswin(M,sigma)

n=(-(M-1)/2:(M-1)/2)';
w = exp(-n.*n./(2*sigma.*sigma));
```

# Gaussian Window and Transform

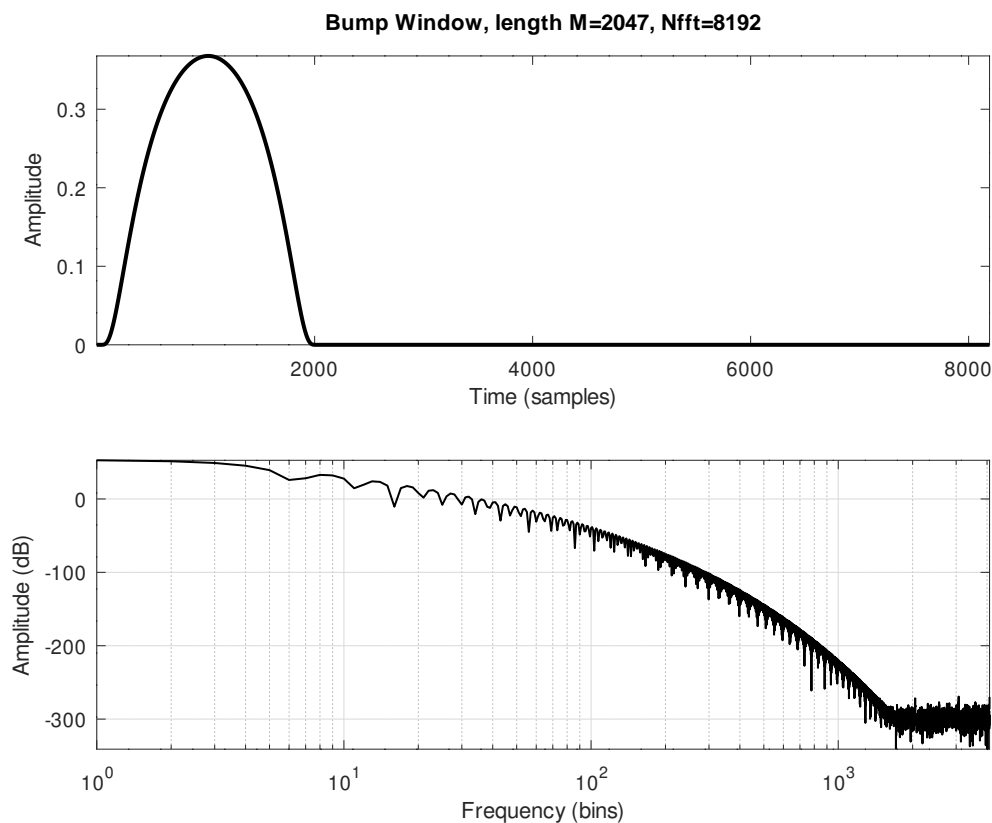


# Bump Window

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$$w(t) = \begin{cases} e^{-\frac{1}{1-t^2}} & |t| < 1, \\ 0 & \text{otherwise} \end{cases}$$

- Infinitely differentiable everywhere (then sample)
- Roll-off rate unbounded (faster than any polynomial)
- Aliasing progressively slows the decay



## More Windows

There is a nice  
collection of window definitions and citations on Wikipedia:  
[https://en.wikipedia.org/wiki/Window\\_function](https://en.wikipedia.org/wiki/Window_function)

# Optimal Windows

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Generally we desire

$$W(\omega) \approx \delta(\omega)$$

- Best results are obtained by formulating this as an *FIR filter design problem*.
- In general, both time-domain and frequency-domain specifications are needed.
- Equivalently, both *magnitude* and *phase* specifications are necessary in the frequency domain.

## Optimal Windows for Audio Coding

Recently, numerically optimized windows have been developed by Dolby which achieve the following objectives:

- Narrow the window in time
- Smooth the onset and decay in time
- Reduce sidelobes below the *worst-case masking threshold*



# Conclusion

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- There is rarely a closed form expression for an optimal window in practice.
- The hardest task is formulating the *ideal error criterion*.
- Given an error criterion, it is usually straightforward to minimize it numerically with respect to the window samples  $w$ .